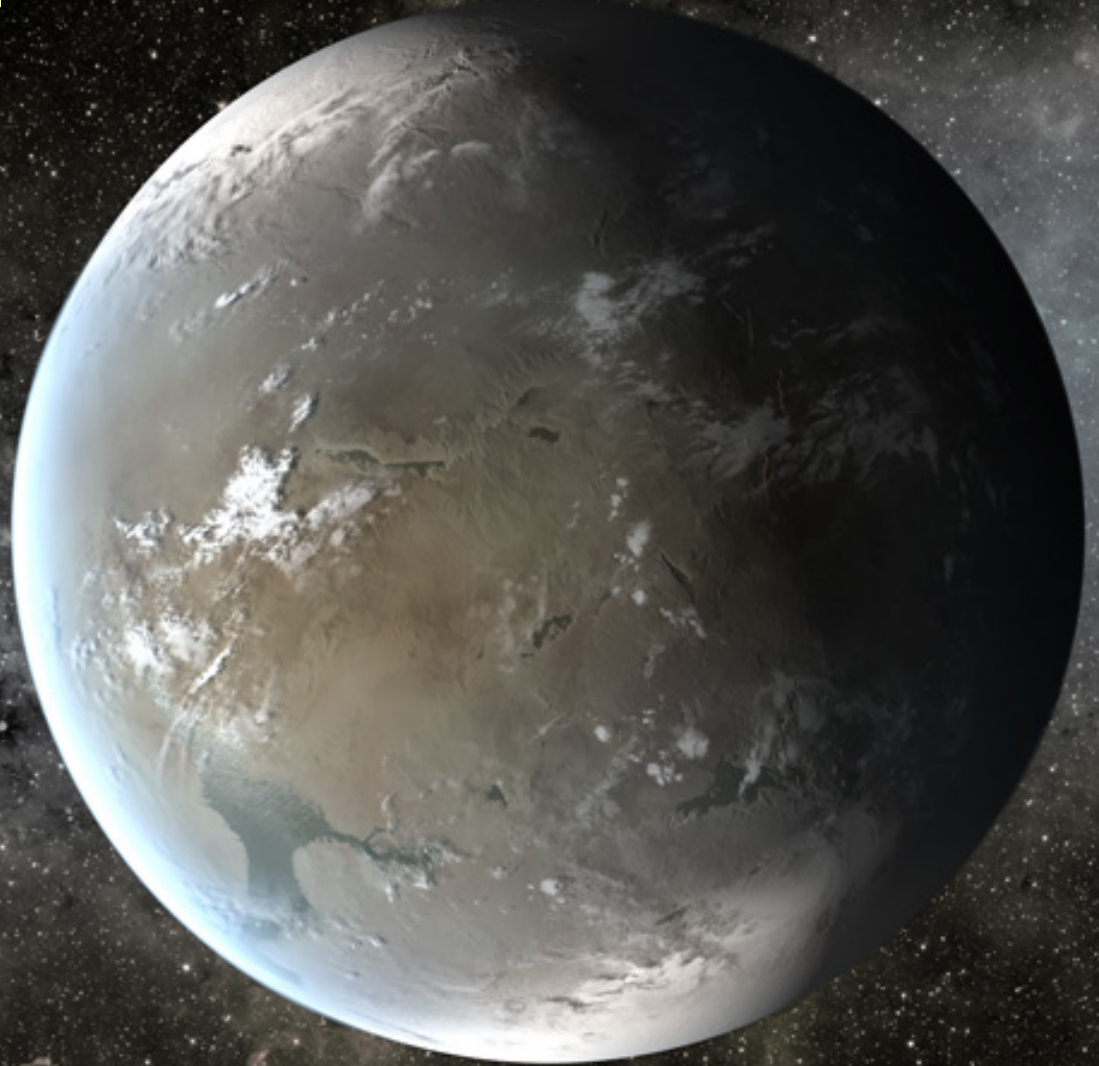


The Search for Other Earths

Class 8: Looking
Forward

Steve Bryson

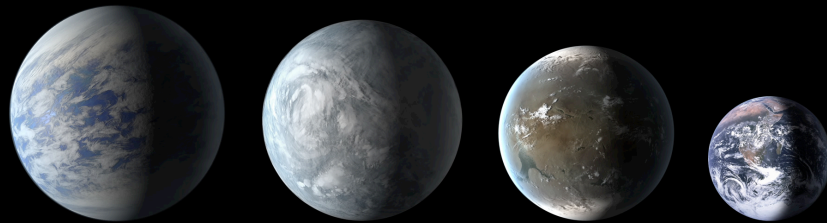


Questions?

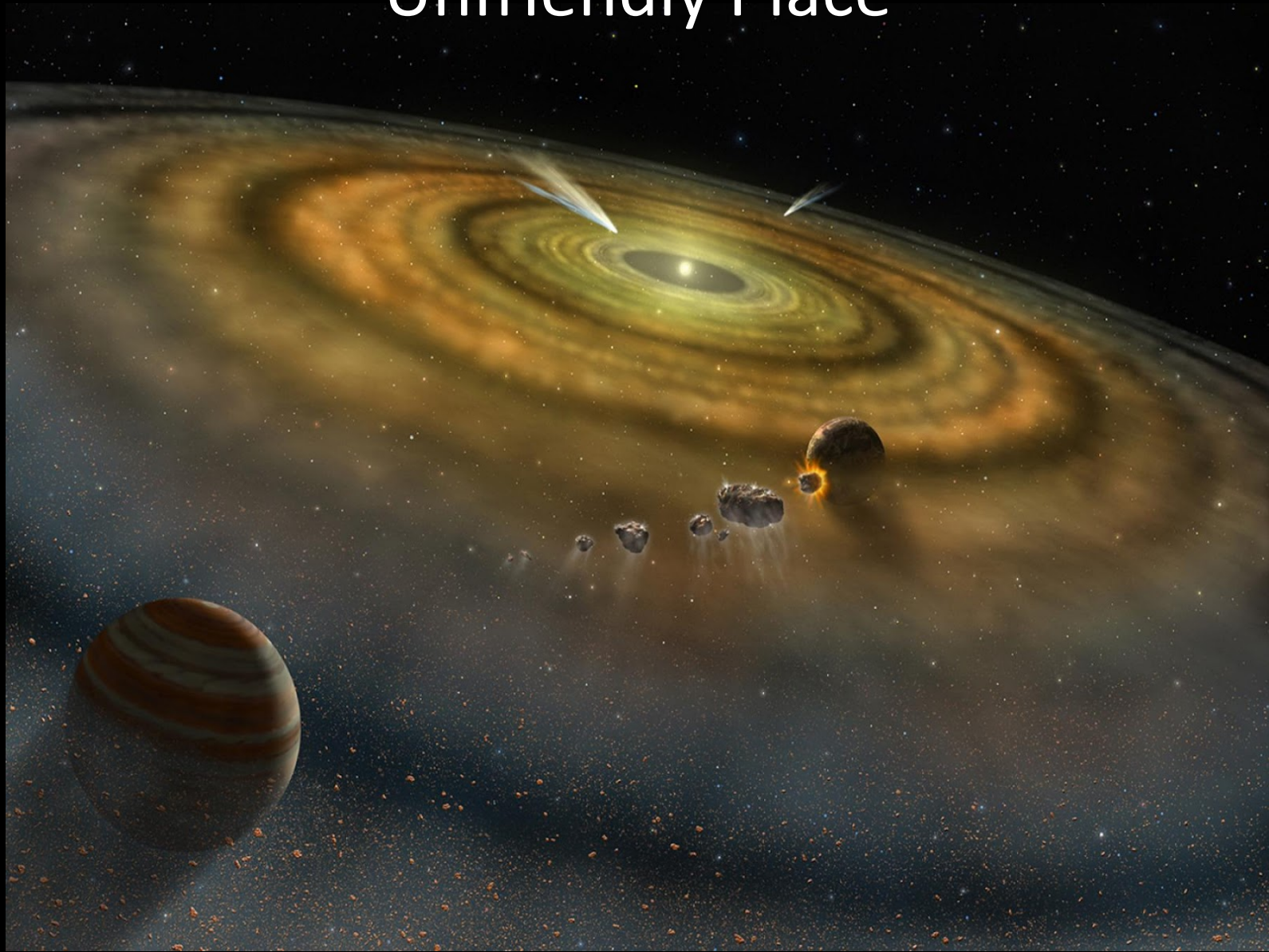


What About Life?

- Currently we can't detect life on exoplanets
- We know a lot about how biology works
- But not much about how biology got started
- So we don't know what is required for life
 - Stick to what we know
 - But don't ignore other possibilities
- Perhaps Super Earths are good for life
 - More active geology provides more good minerals

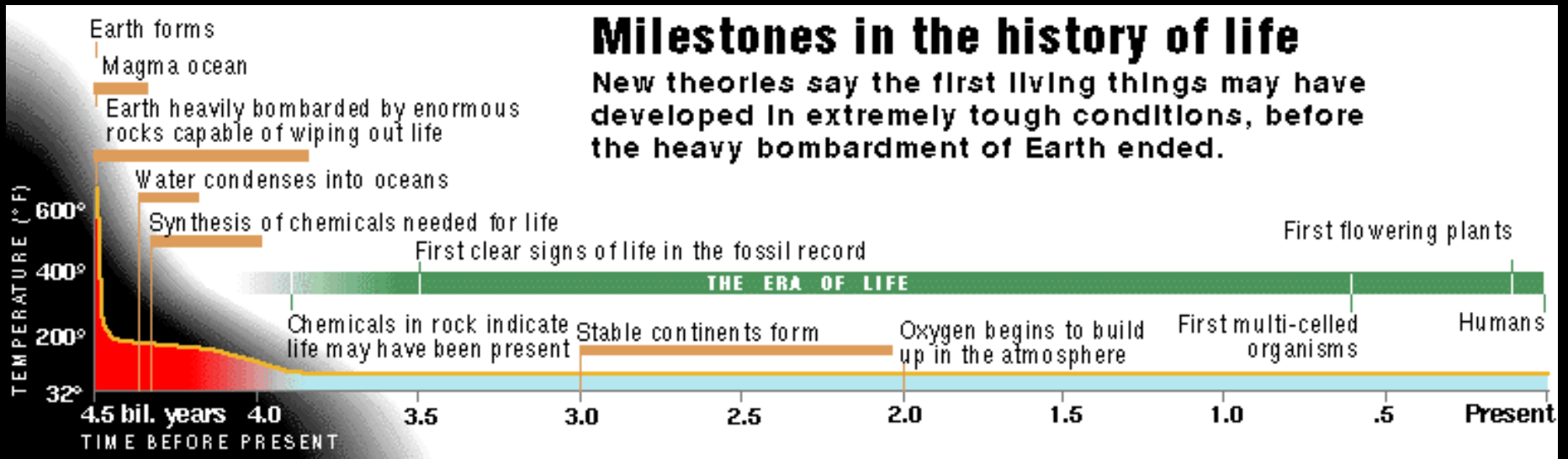


The Infant Solar System was a Very Unfriendly Place



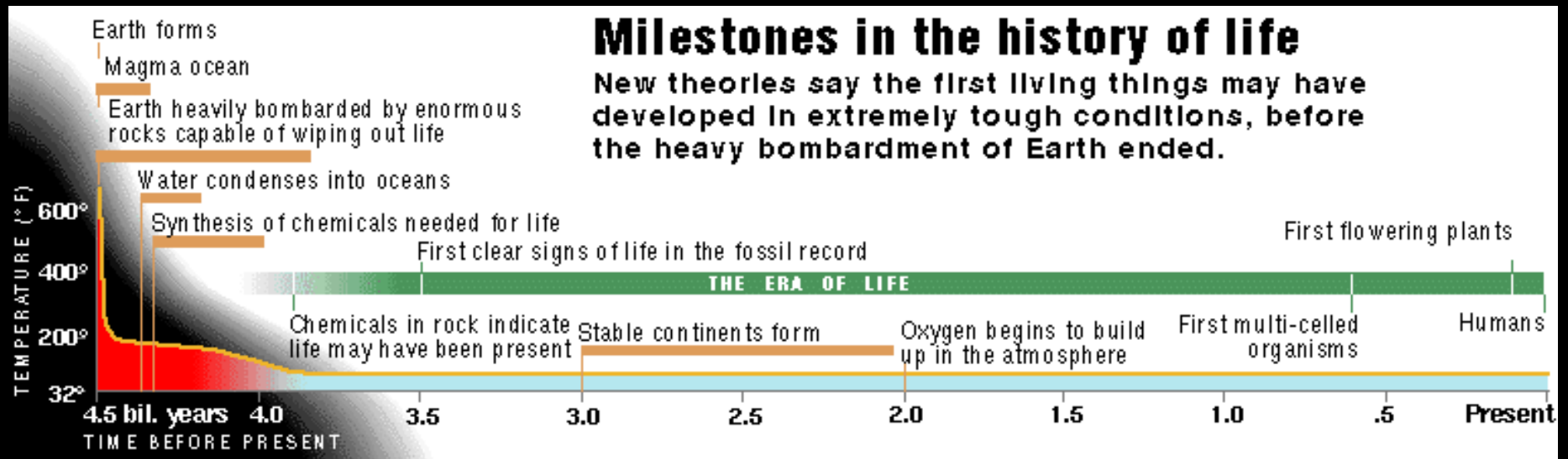
Life on Earth Happened Quickly

- We have evidence for simple life almost as soon as Earth was cool enough to allow it
 - Does this mean that the creation of life is probable?



But Complex Life Took a Long Time

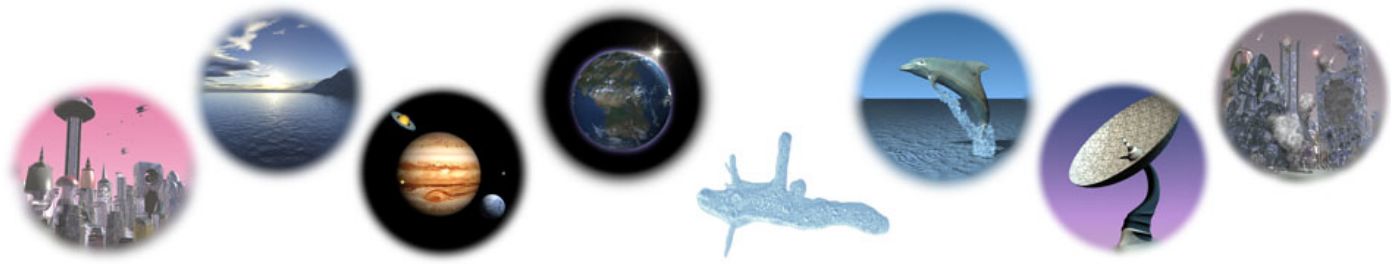
- Multi-celled animals took a lot longer
 - Does this mean that the evolution of complex life is less probable?



What About Other Civilizations?

- Depends on
 - How many stars
 - What fraction of those stars have planets
 - What fraction of those planets are Earth-size
 - What fraction of those planets have life
 - What fraction of that life evolves intelligence
 - What fraction develop the ability to communicate
 - How long such civilizations live
- Multiply all these together and get the *Drake Equation*

The Drake Equation



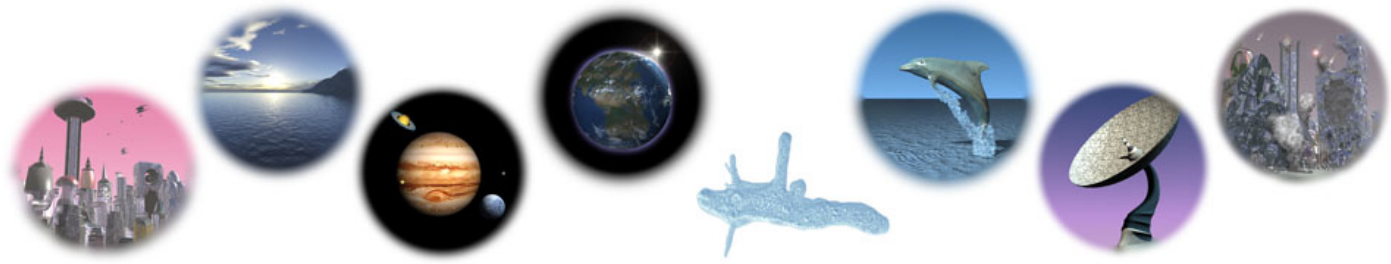
*Drake
Equation*

$$N = R_* \times f_p \times n_e \times f_l \times f_i \times f_c \times L$$

Graphic by Chris Short
<http://www.cyrstudio.com/short>

- 30 years ago we only knew how many stars there are (R_*)

The Drake Equation



*Drake
Equation*

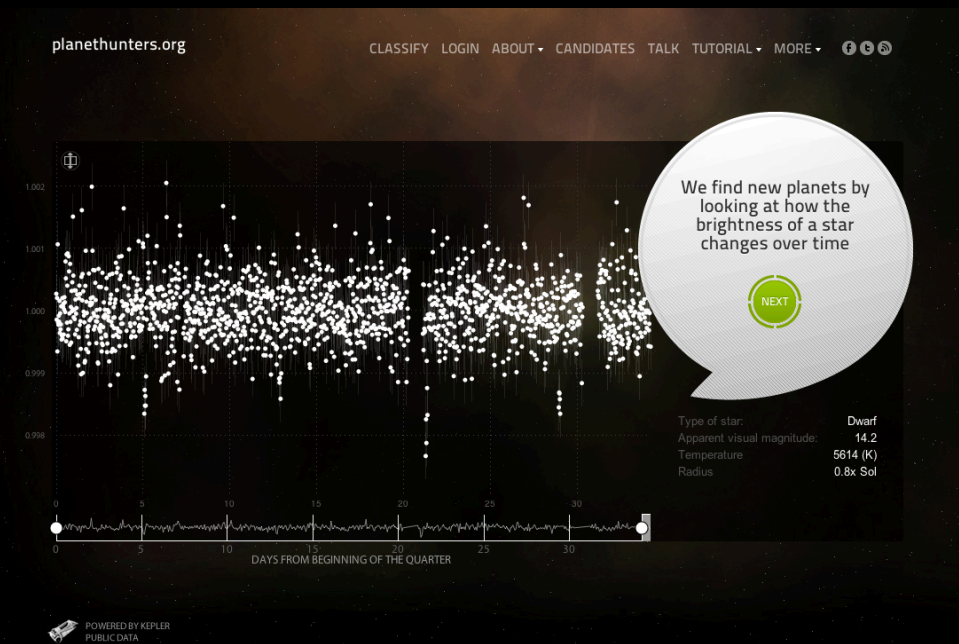
$$N = R_* \times f_p \times n_e \times f_l \times f_i \times f_c \times L$$

Graphic by Chris Short
<http://www.cyrstudio.com/short>

- Now we know what fraction of stars have planets (all) and the fraction of planets that are Earth-size (~15%)
- We have to dig more deeply in the data to measure how many Earth-like planets are in the habitable zone

Continuing to Dig Into the Data

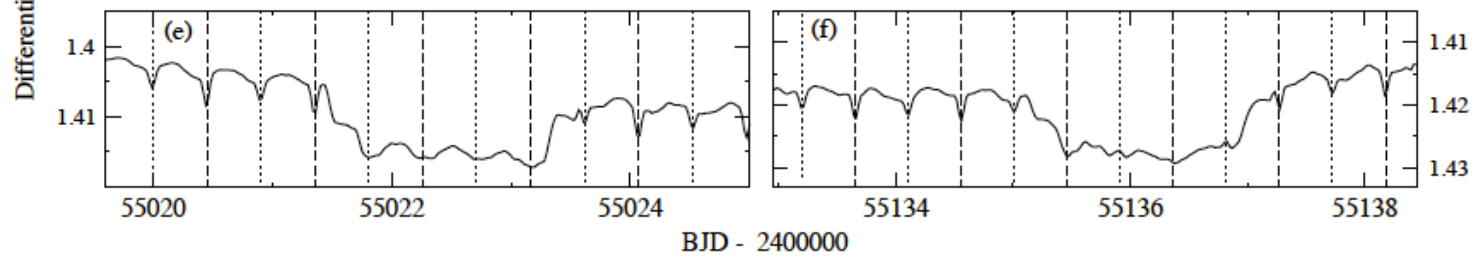
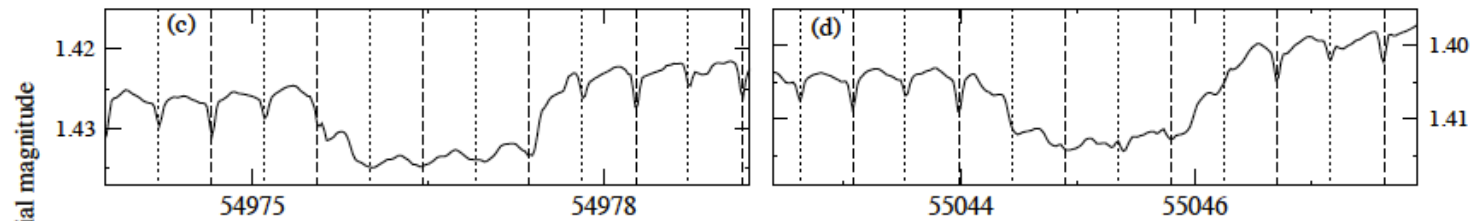
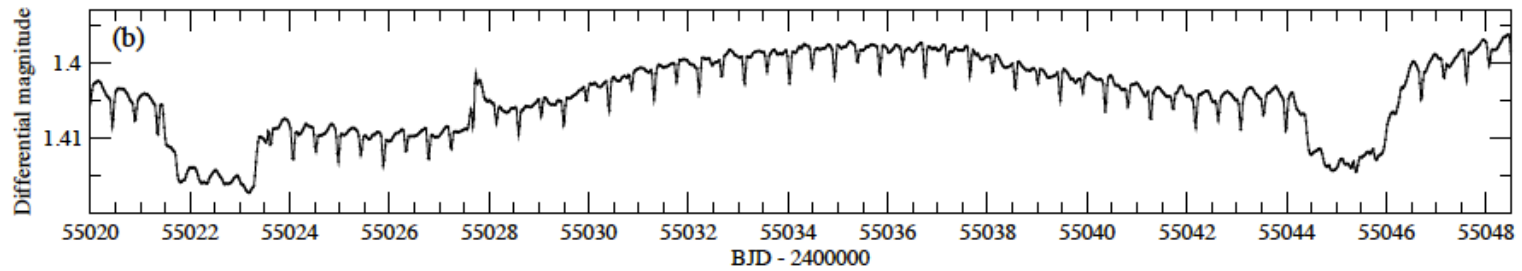
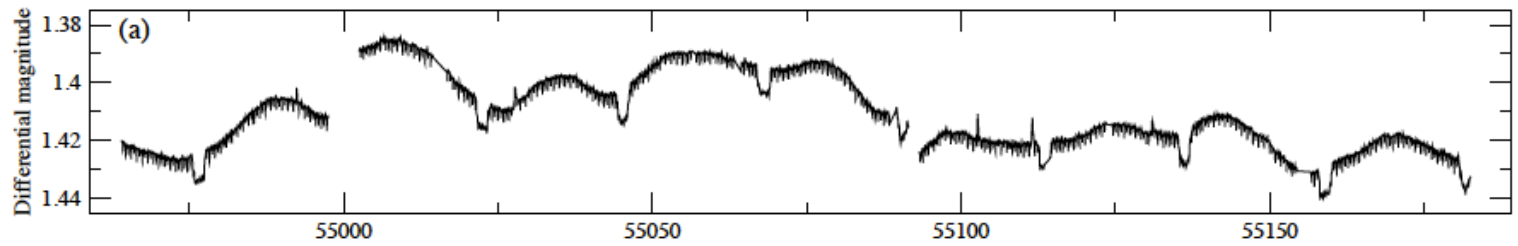
- Astronomers will continue to look at Kepler data for years to come
- You can help! www.planethunters.org
 - Help find new planets!
 - Some already found
 - Missed by the Kepler team



Kepler Also Observes Stars

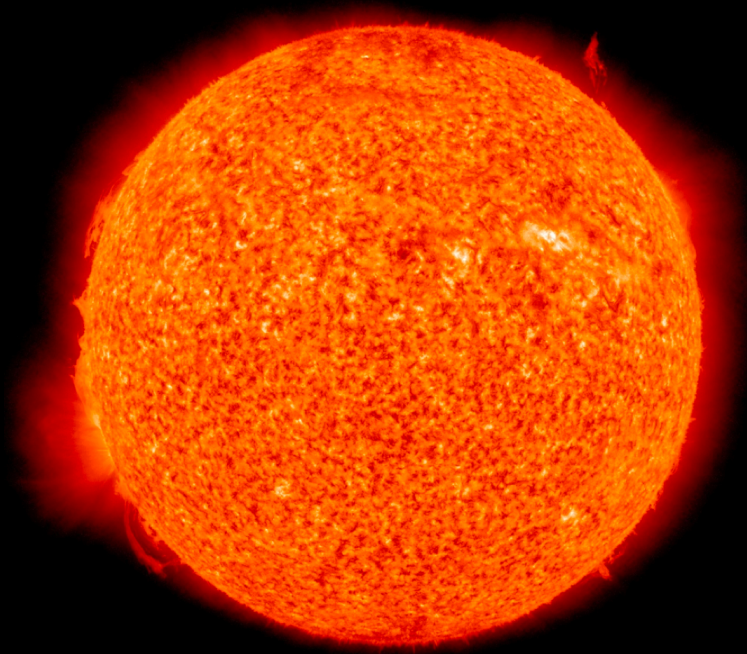
- Most Boring Telescope?
 - It always looks at the same thing
- But wait! Kepler always looks at the same thing!
 - Fantastic opportunity to make long-term observations of stars at high precision

“Trinity”



Trinity

A



12.4 R_☉

HD 181068

B



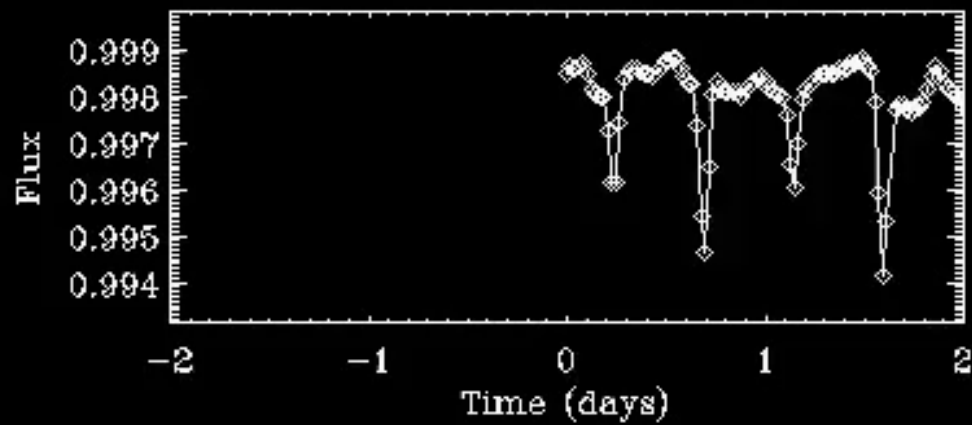
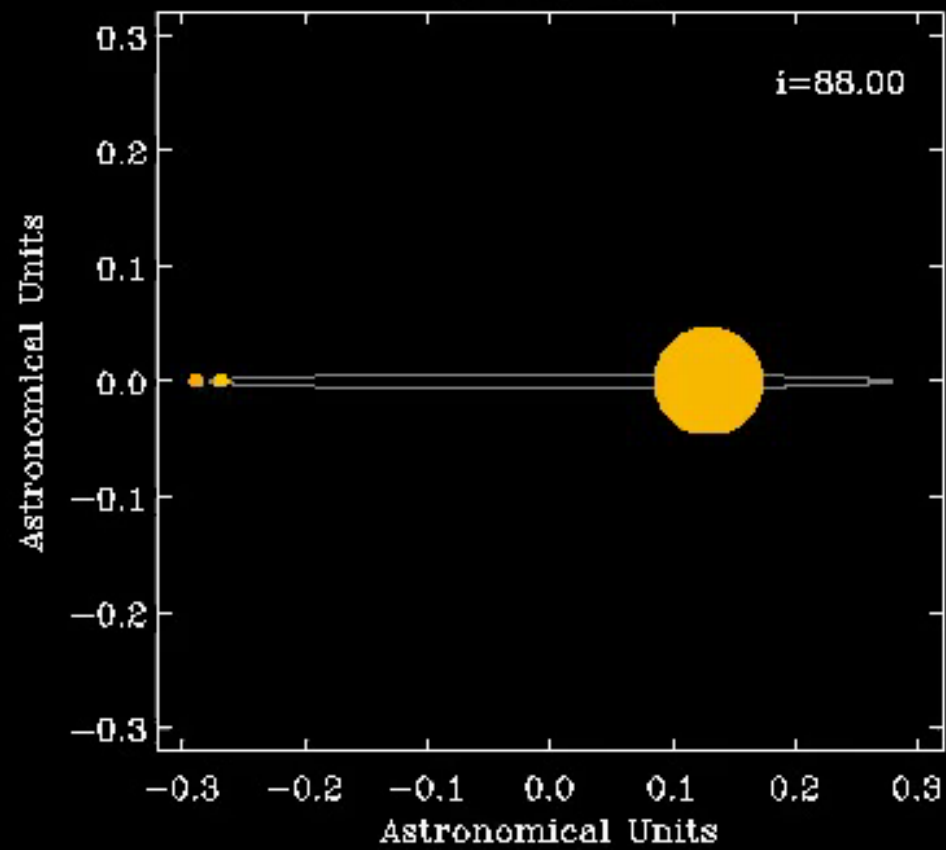
0.8 R_☉

C



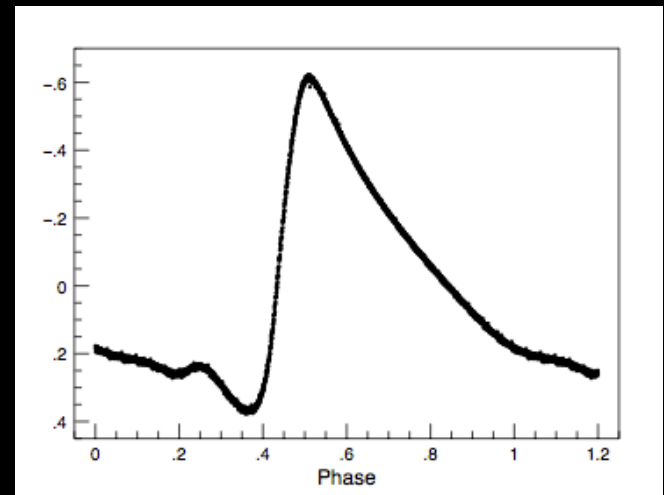
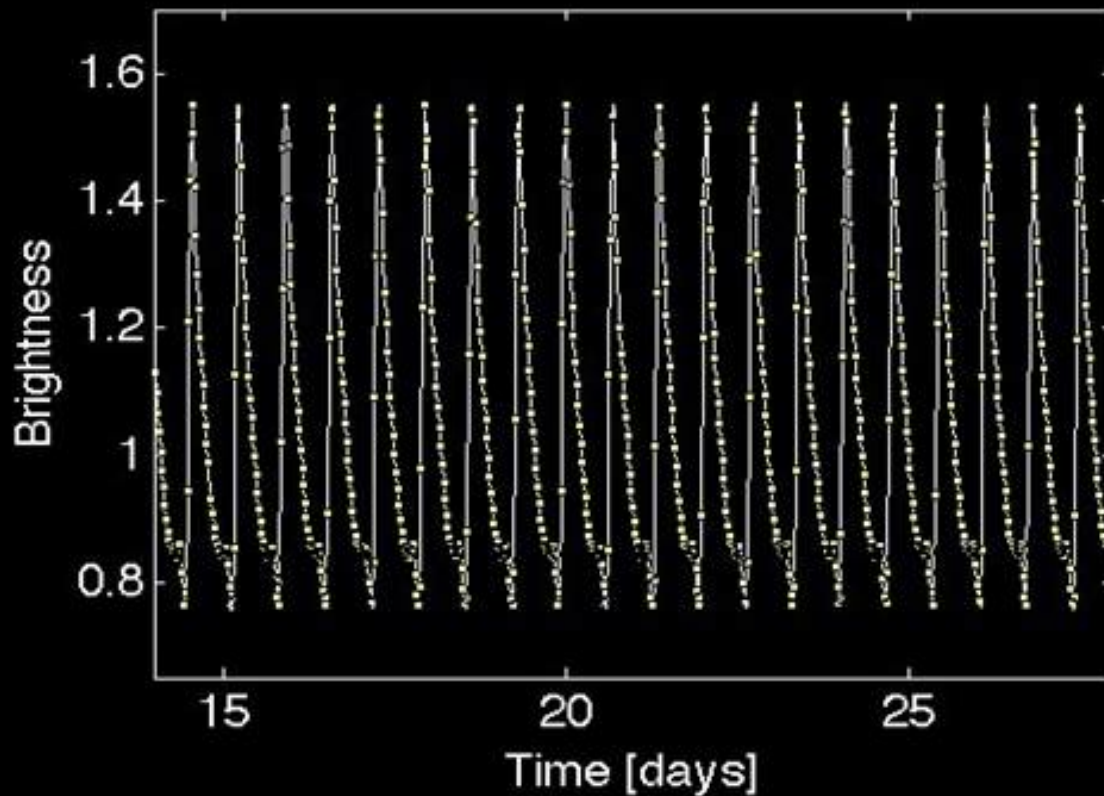
0.7 R_☉

HD181068



RR Lyrae Stars

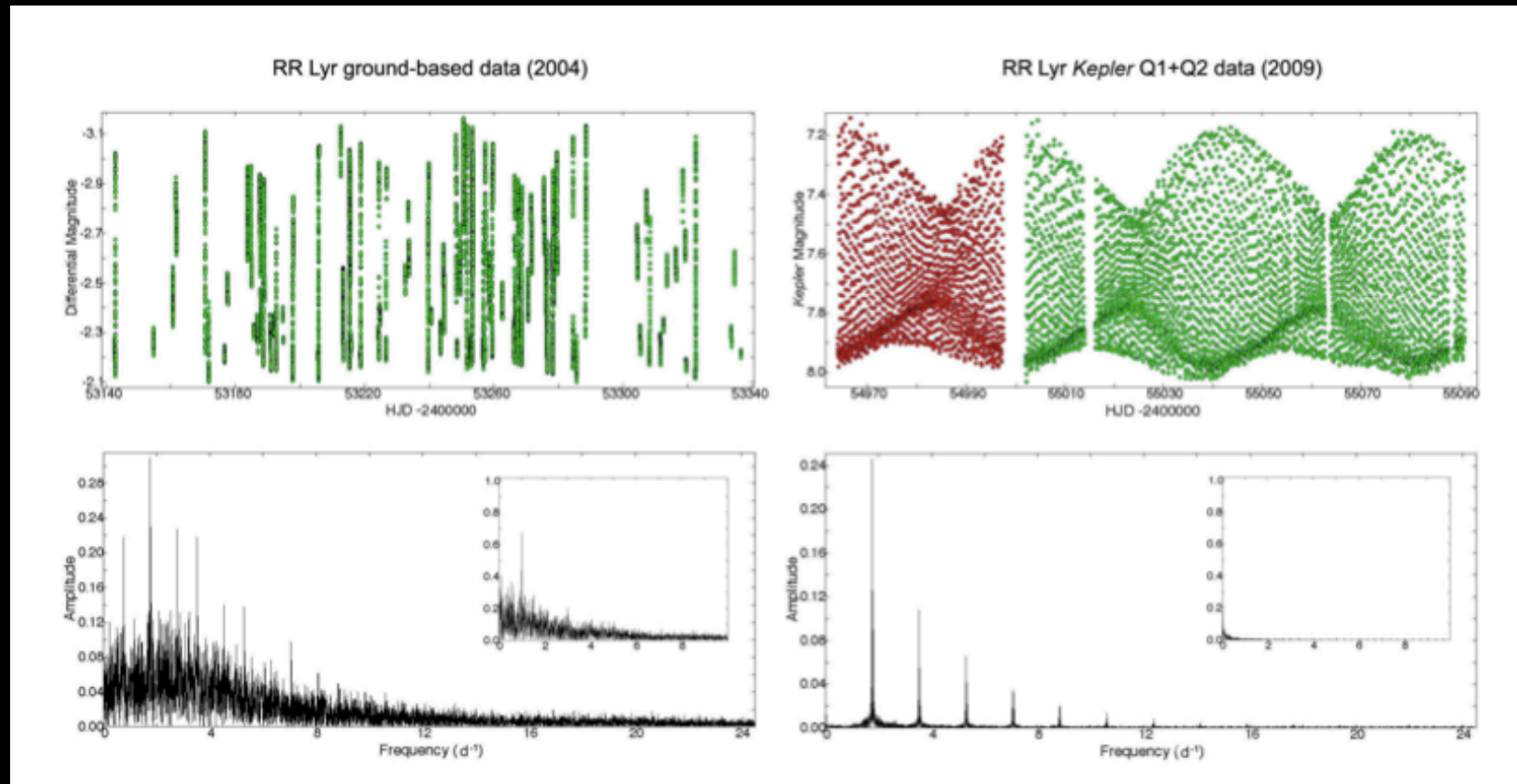
- Variable stars that double in brightness in only a couple hours more than once a day



One period close up

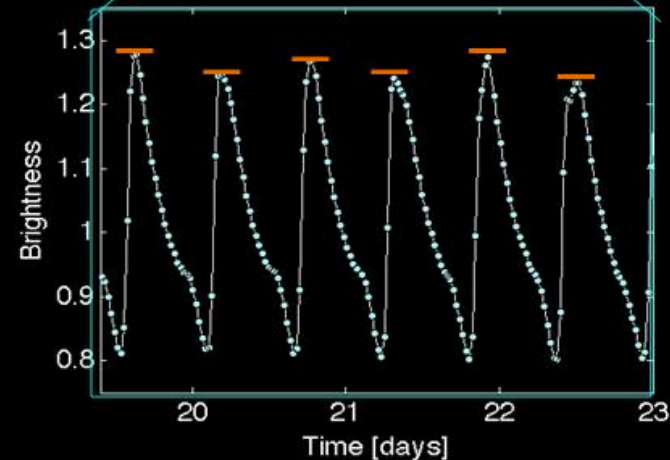
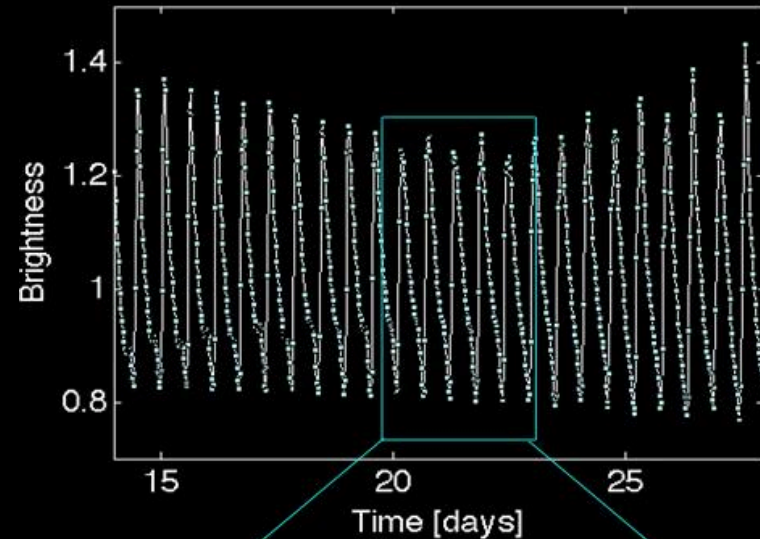
Example of RR Lyrae

- Happens to be in the Kepler field
 - New discovery of period doubling

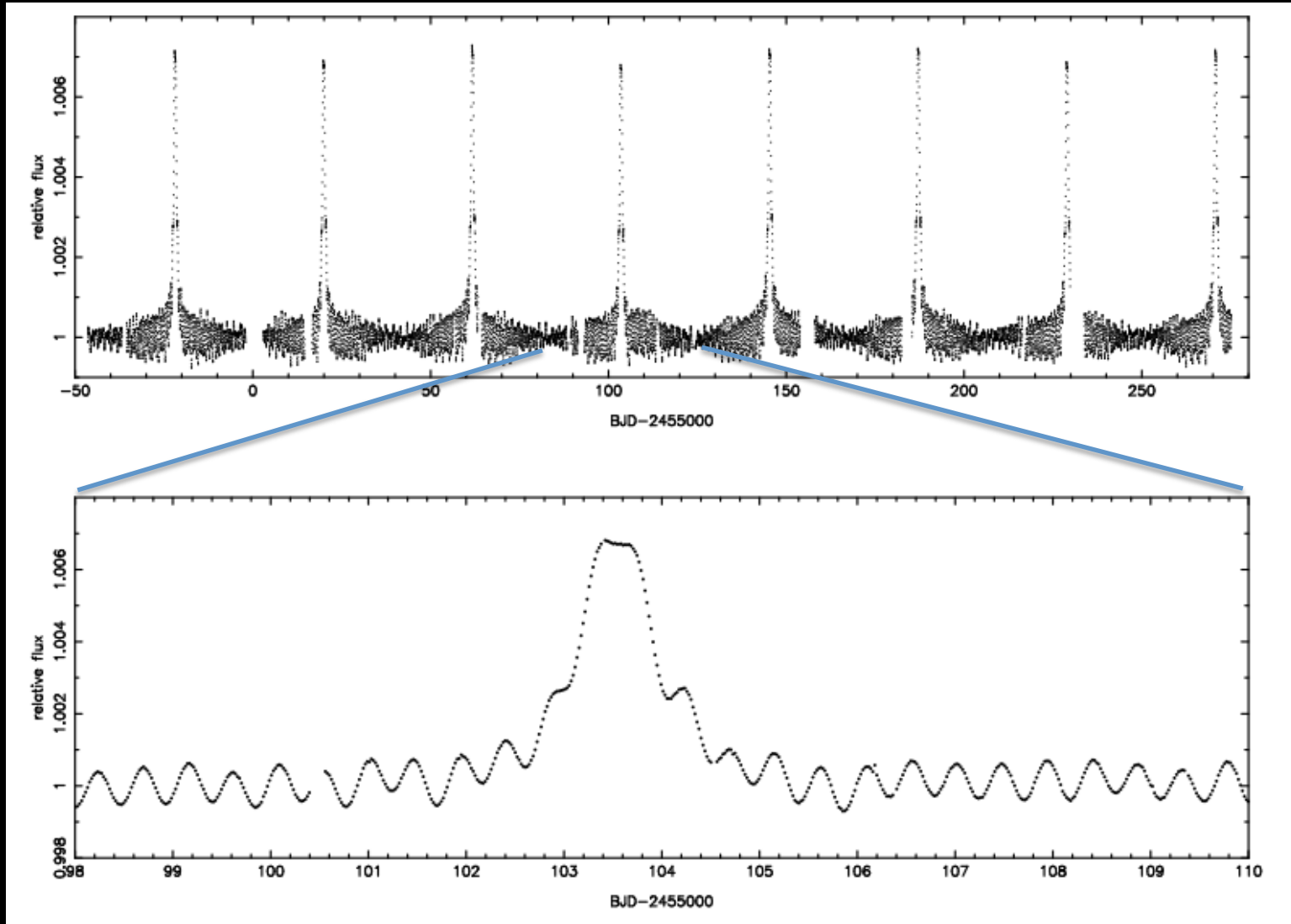


RR Lyrae Stars

- On some RR Lyrae stars the variation is very steady, like on the previous slide
 - We have a good idea why these steady variations happen
- On other RR Lyrae stars the variation is complicated
- We don't understand why there are these complications, but Kepler is helping make better theories

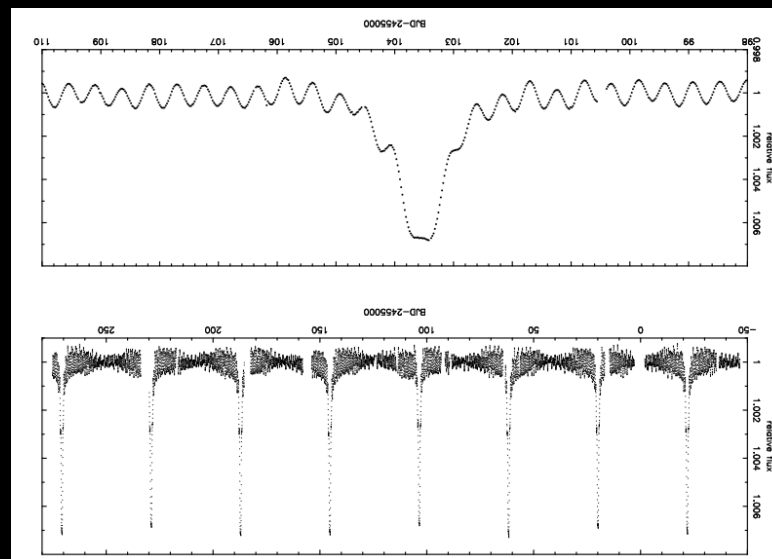


The Mystery of KOI-54



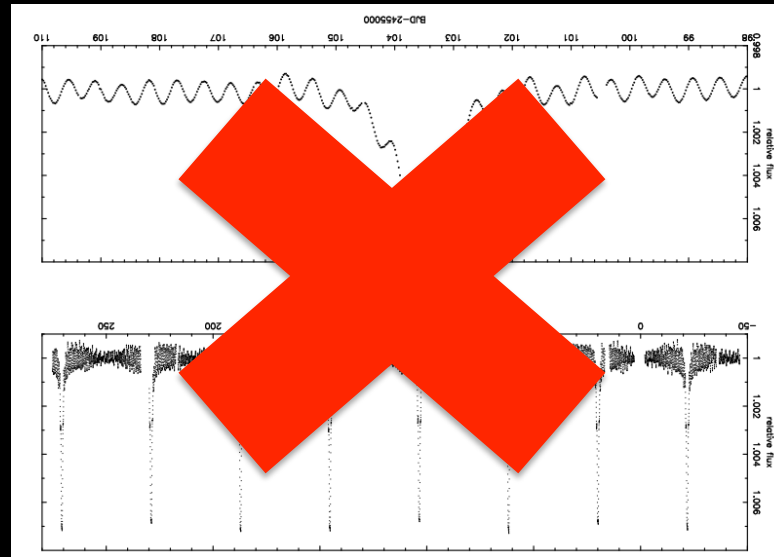
KOI-54: What Is It?

- Upside down?



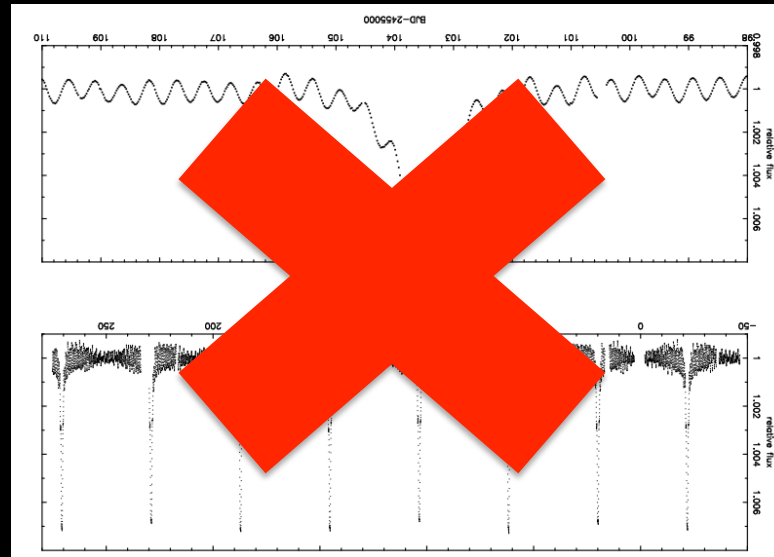
KOI-54: What Is It?

- Upside down?

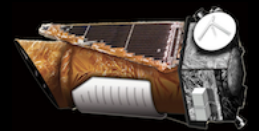
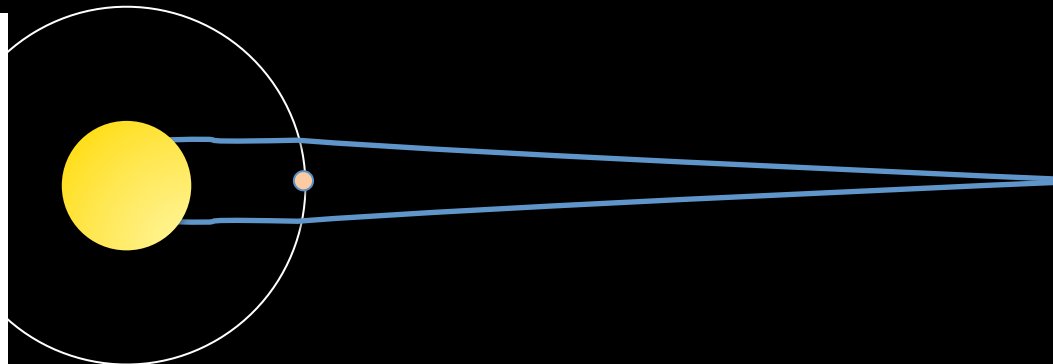
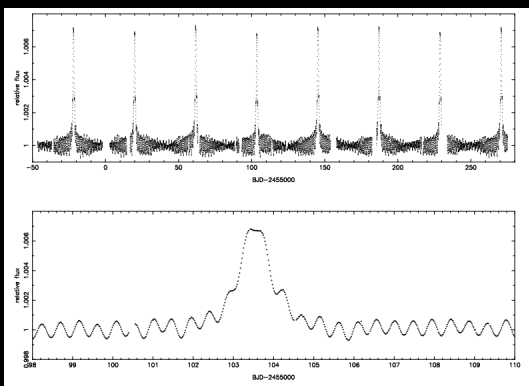


KOI-54: What Is It?

- Upside down?

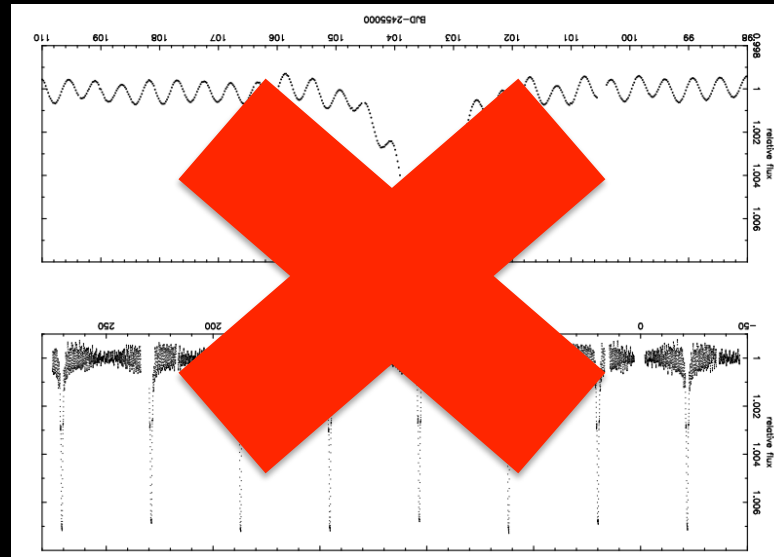


- Gravitational Lensing by a Neutron Star or Black Hole?

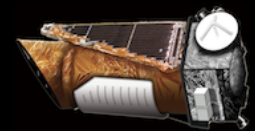
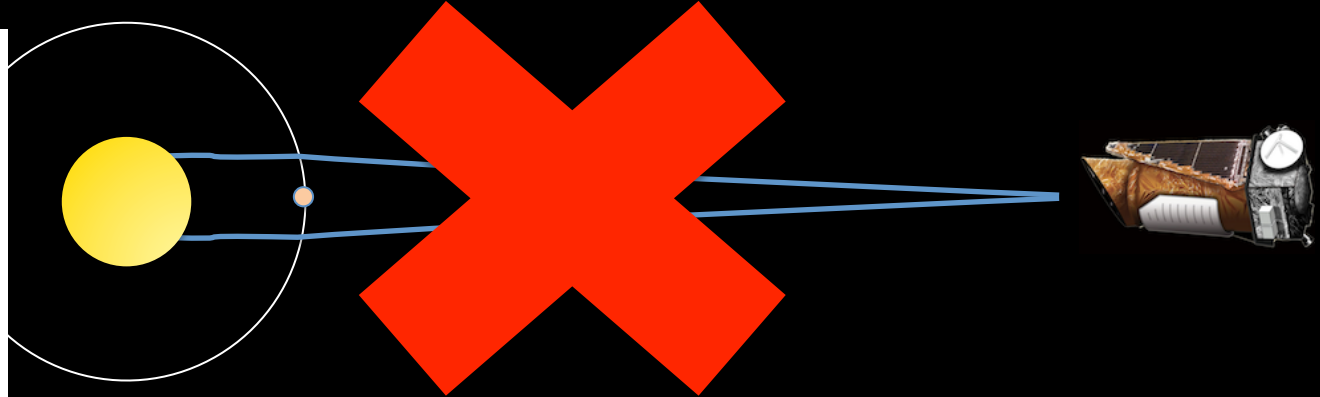
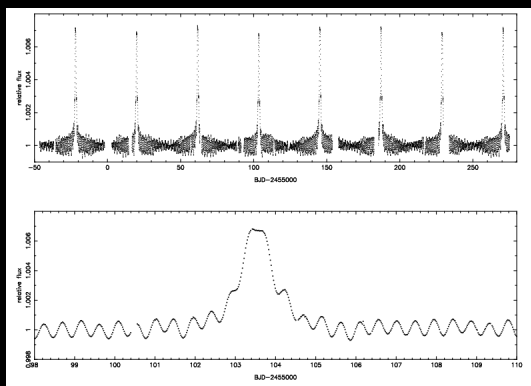


KOI-54: What Is It?

- Upside down?

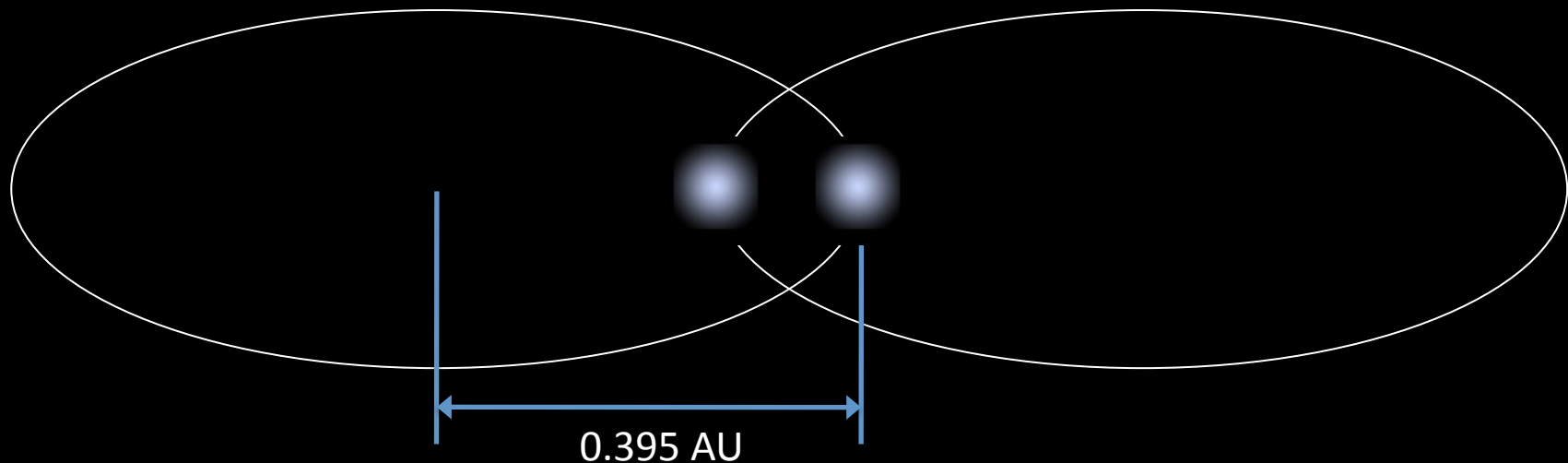


- Gravitational Lensing by a Neutron Star or Black Hole?

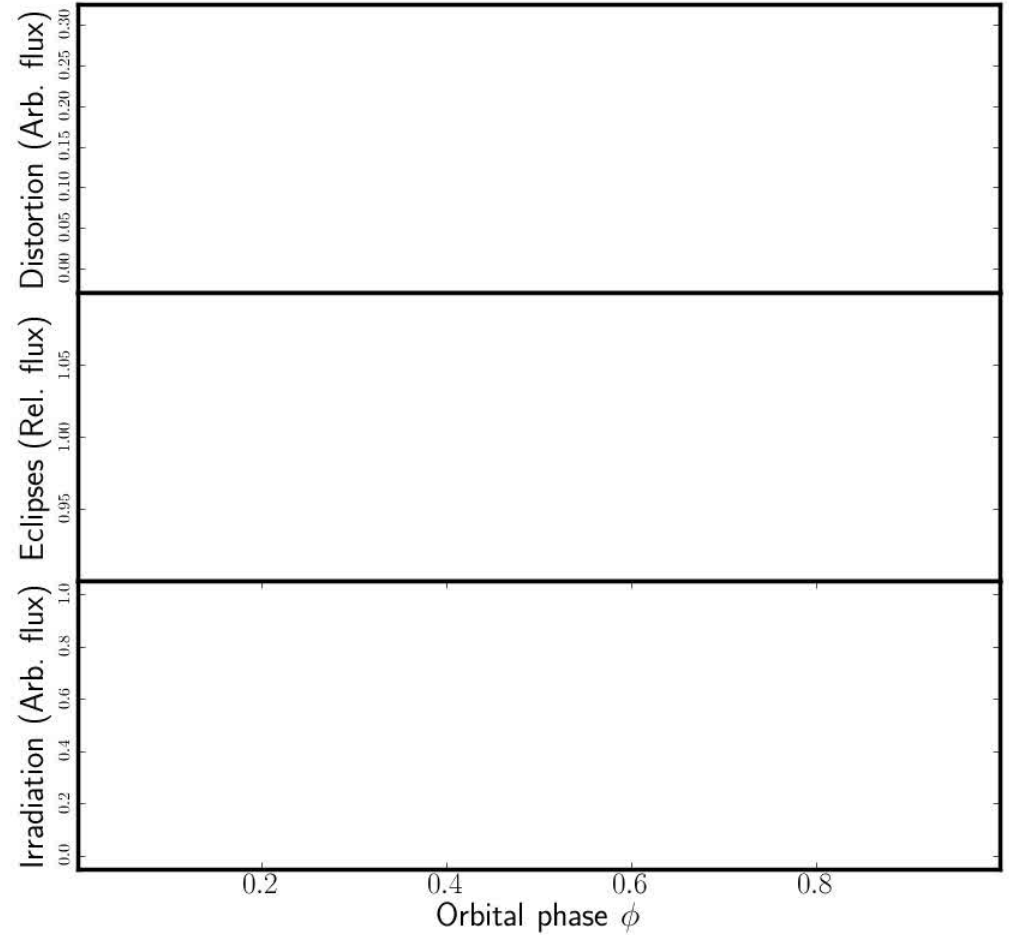
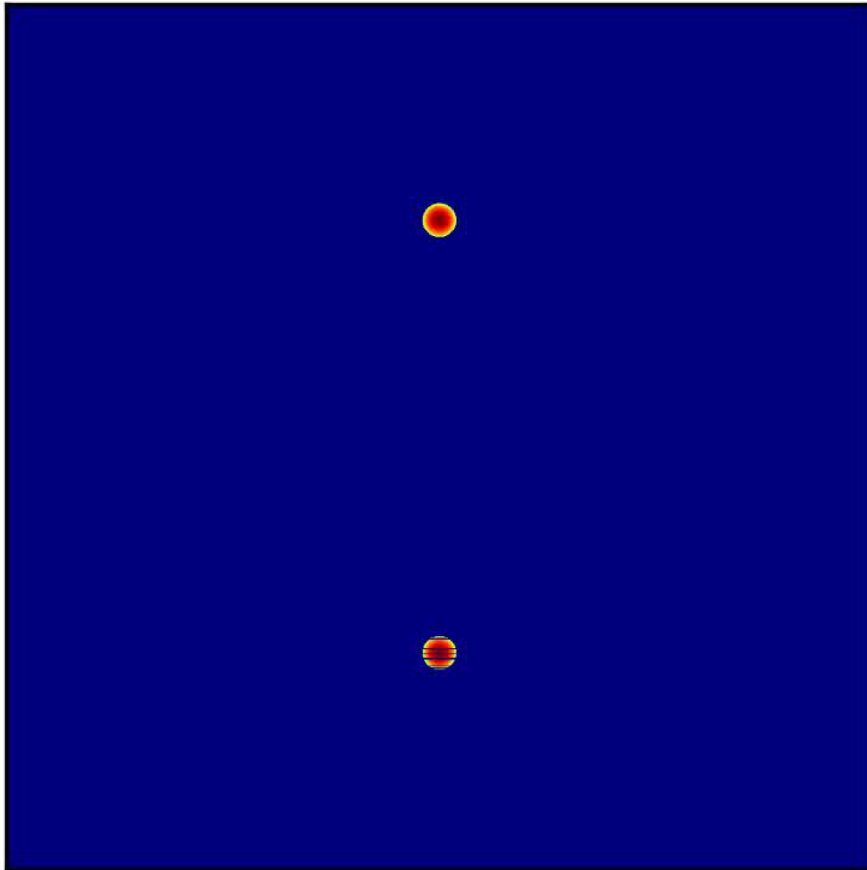


KOI-54: The Answer

- Two, almost identical A (blue giant) stars in very elliptical orbits (41.8 day period)
 - Every 41.8 days the stars are very close together and heat each other up increasing the brightness
 - The close encounter also creates oscillations in both star's size

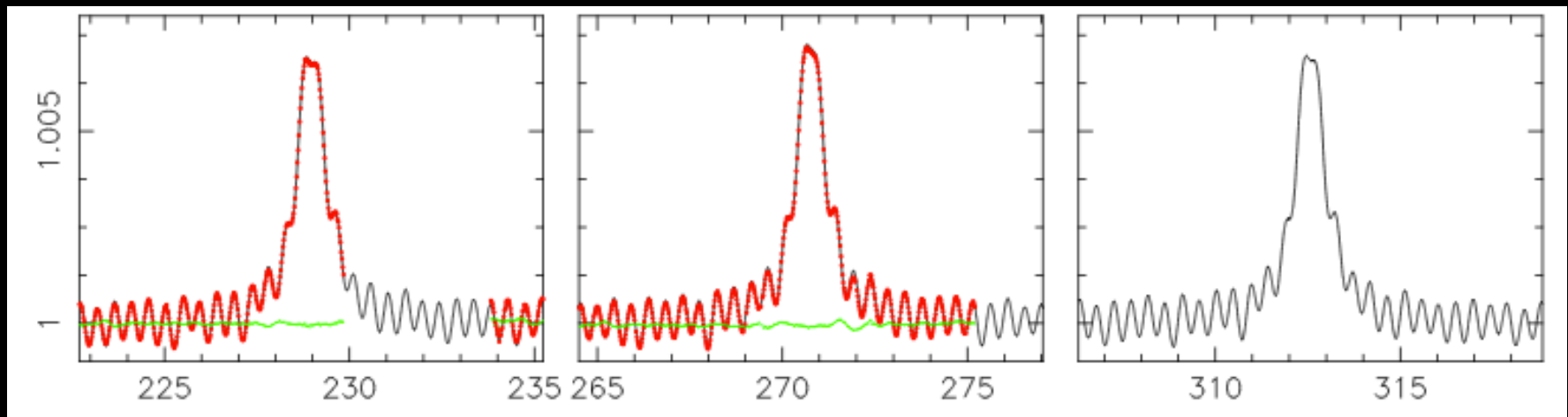


KOI-54: The Answer



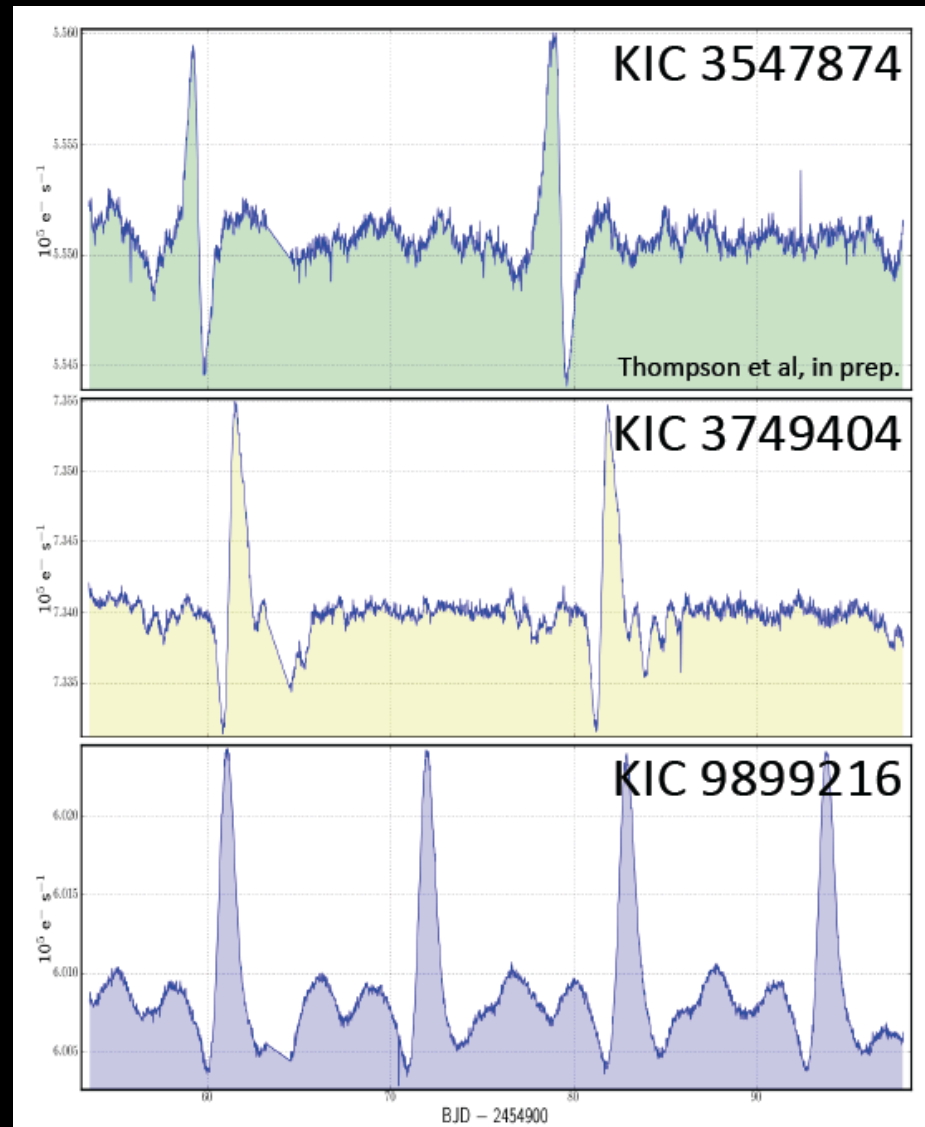
KOI-54: How Do We Know?

- We make a theoretical model of how such stars would behave if it were true
- The data matches the model exactly!



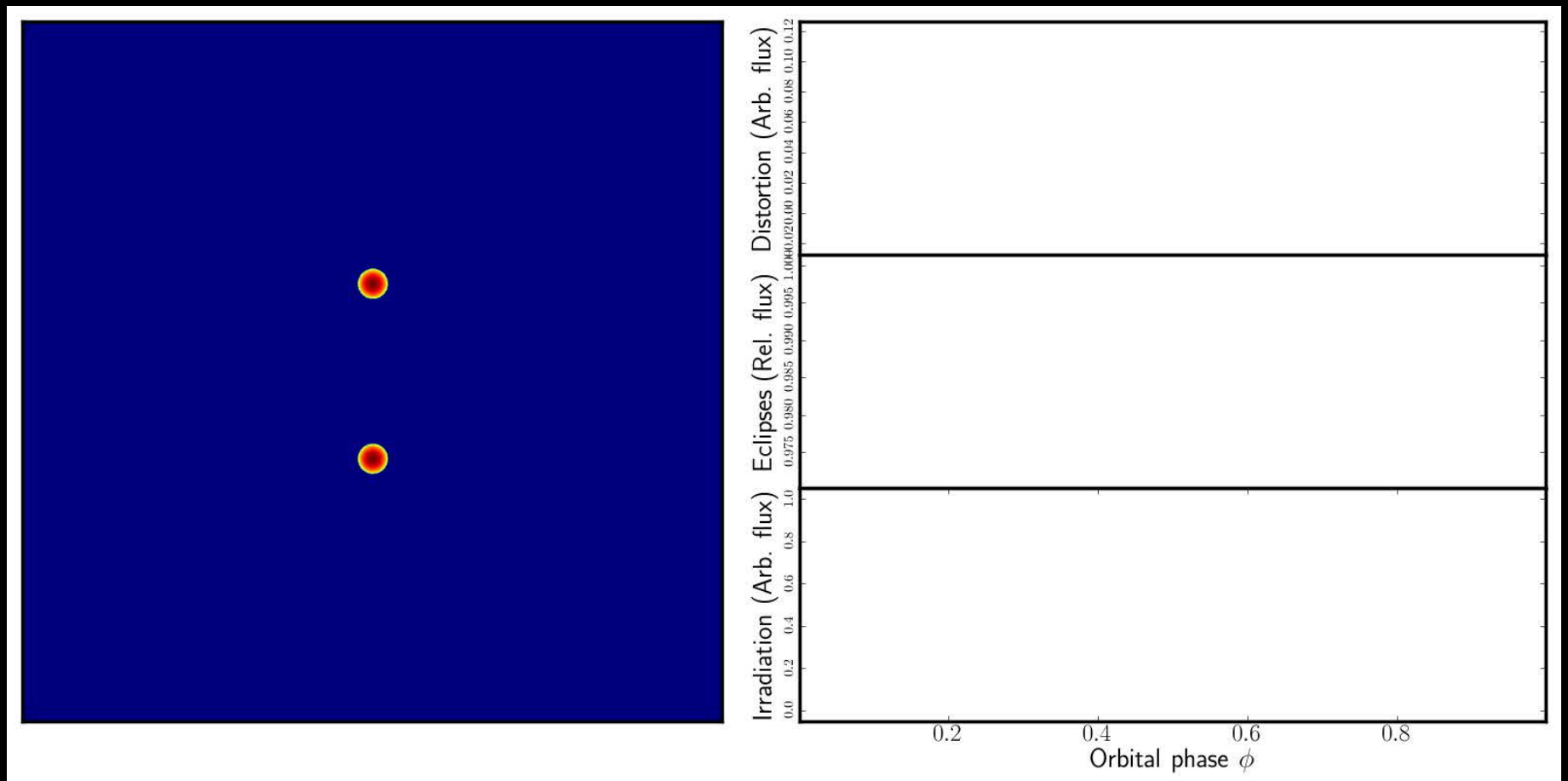
Red: observed light curve
Gray: model

Heartbeat Stars



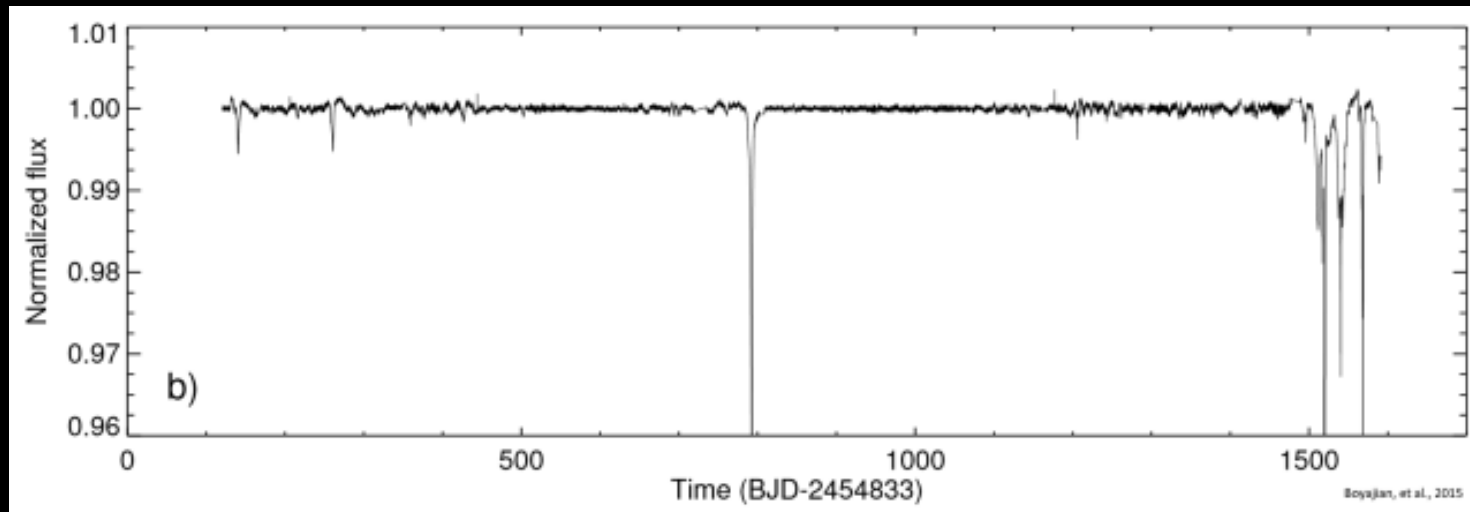
Heartbeat Stars are Close Binaries

- Like KOI-54 but also eclipsing!

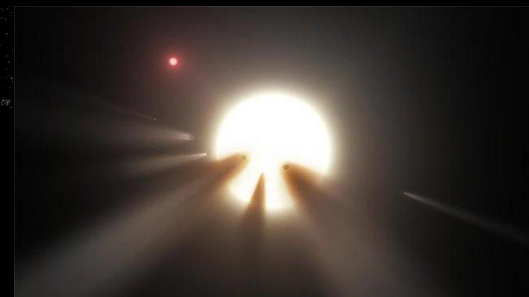
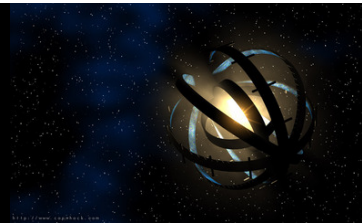


Mysterious KIC 8462852

- Very strange light curve



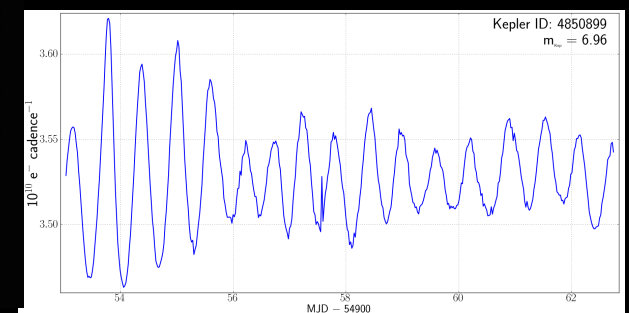
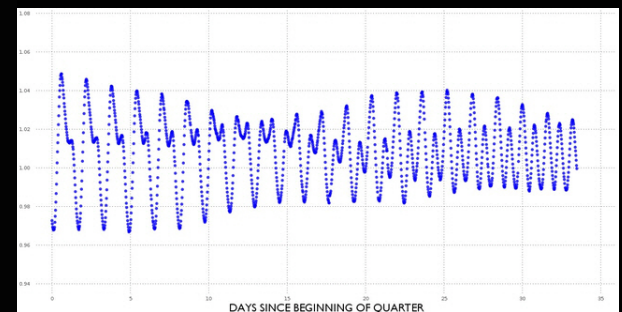
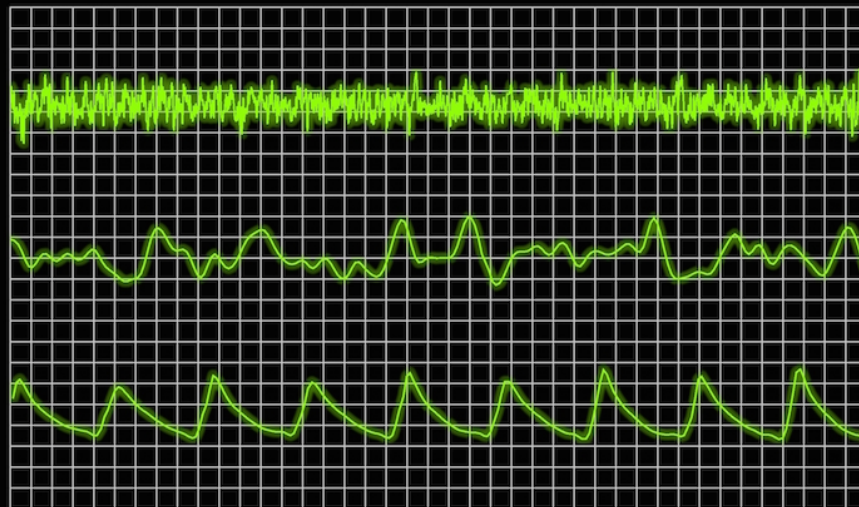
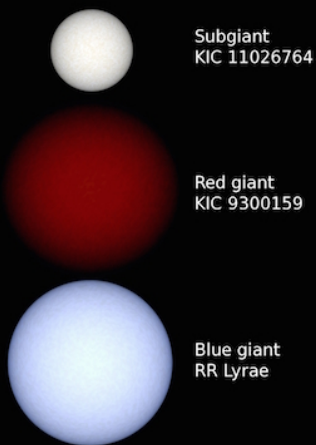
- Alien Megastructures?
 - Almost certainly not
- Comet-like swarm?
 - More likely



Kepler Measures Starlight!

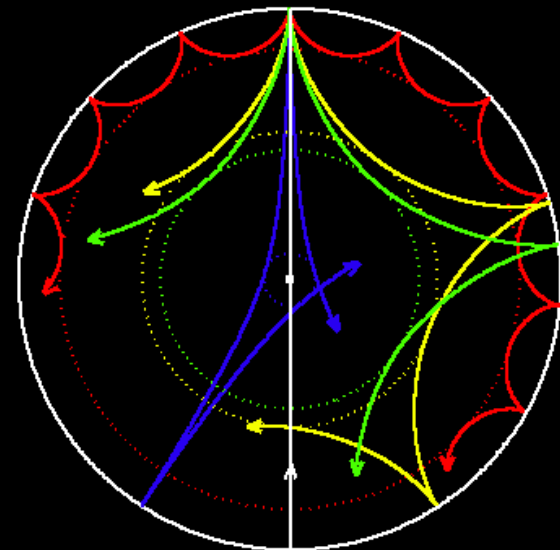
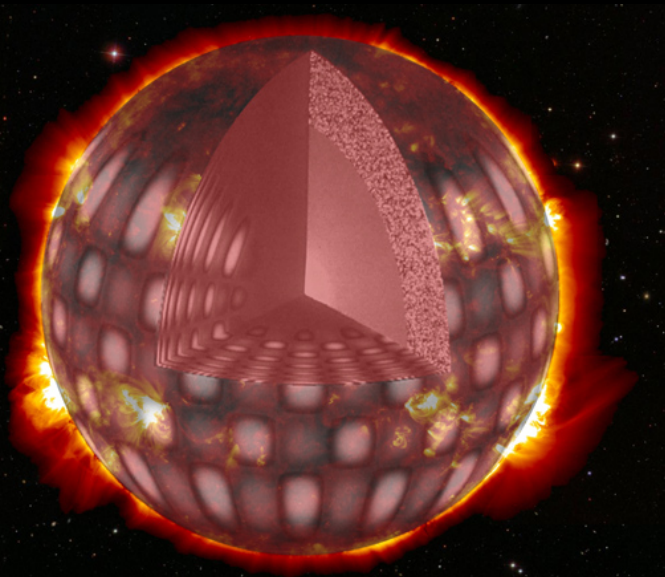
- Changes in the starlight tell us almost everything about the star
 - Size, temperature, structure, age...

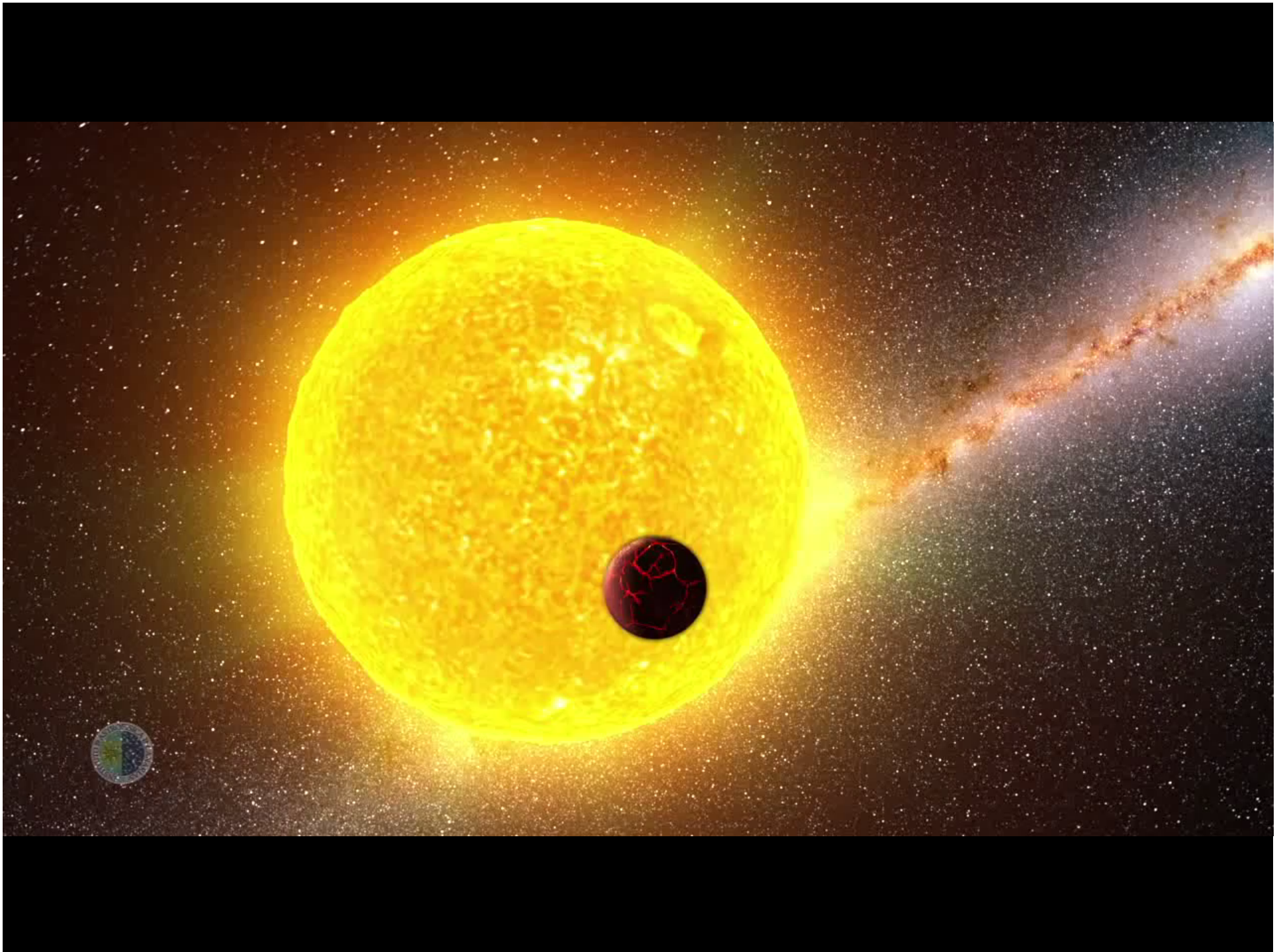
TAKING THE "PULSE" OF STARS



Asteroseismology

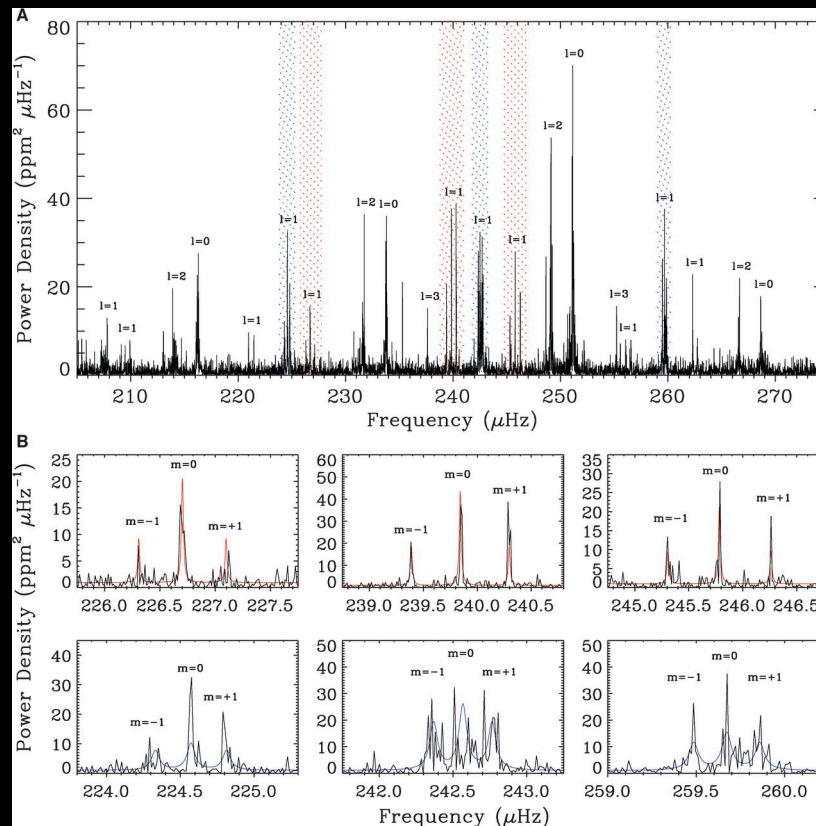
- Subtle oscillations in a star's brightness reflect acoustic waves bouncing around inside the star
- These waves tell us the temperature and density structure of the star
 - Just like in geological seismology





Kepler's Asteroseismology

- The precision and long time base to find planets is perfect for asteroseismology!



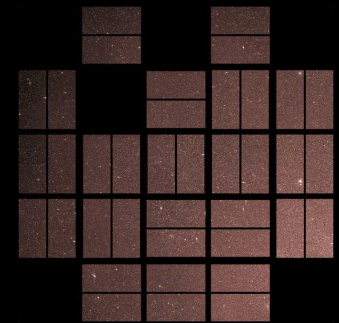
Exploding Stars in Other Galaxies

- Kepler caught the initial flash of an exploding star
 - Predicted but never before seen



And Then Kepler Broke

- Lost the ability to control pointing in 3 dimensions
 - Can control two, such as pitch and yaw
 - Kepler can no longer resist the push of photon pressure from the Sun
- Almost 4 years after launch
- Just after getting a 4-year mission extension
- We also lost almost 10% of the pixels
- But otherwise a wonderful telescope
 - On a wobbly mount

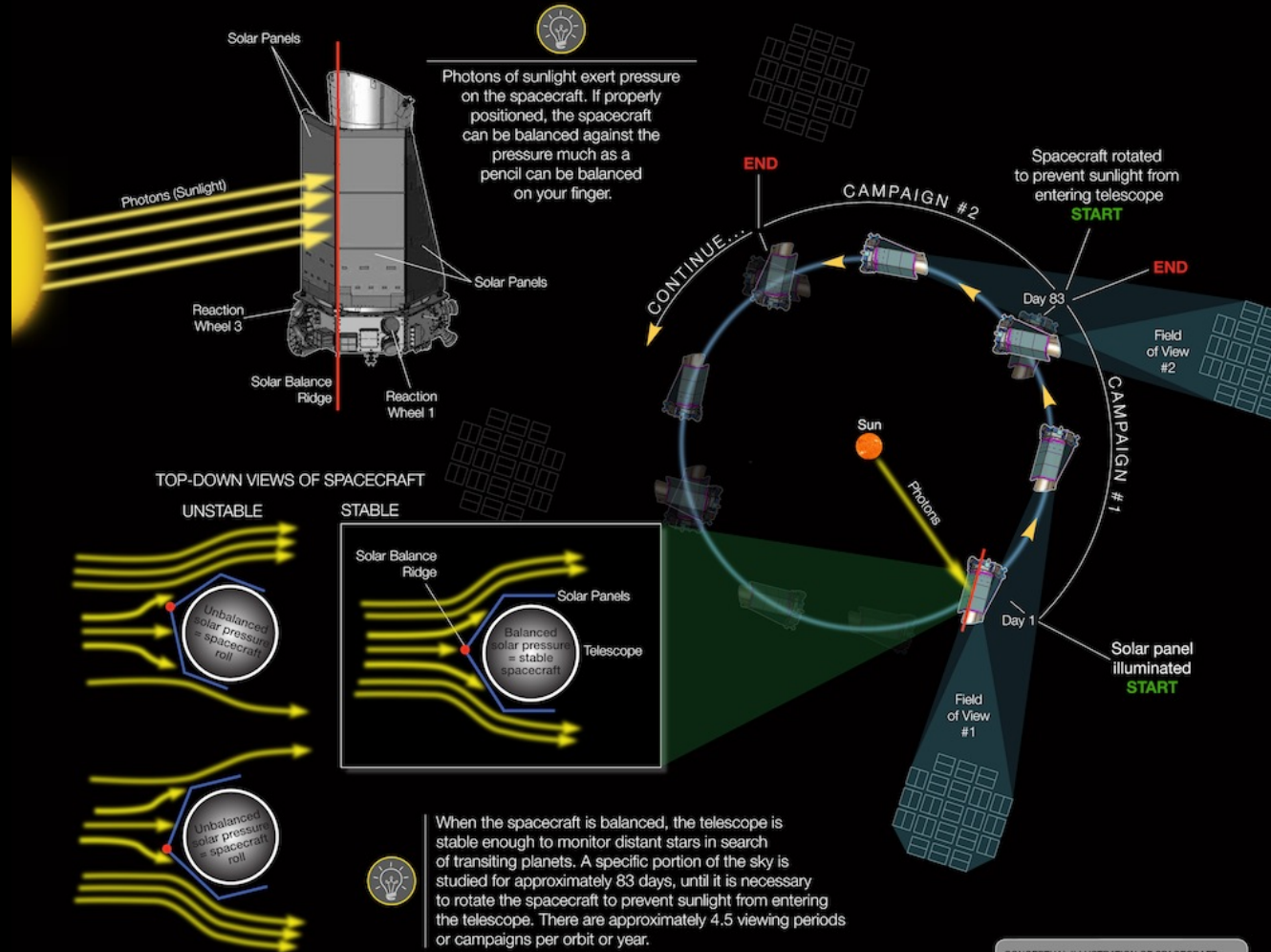


Balance the Photon Pressure to Control Roll!

Kepler's Second Light: How K2 Will Work



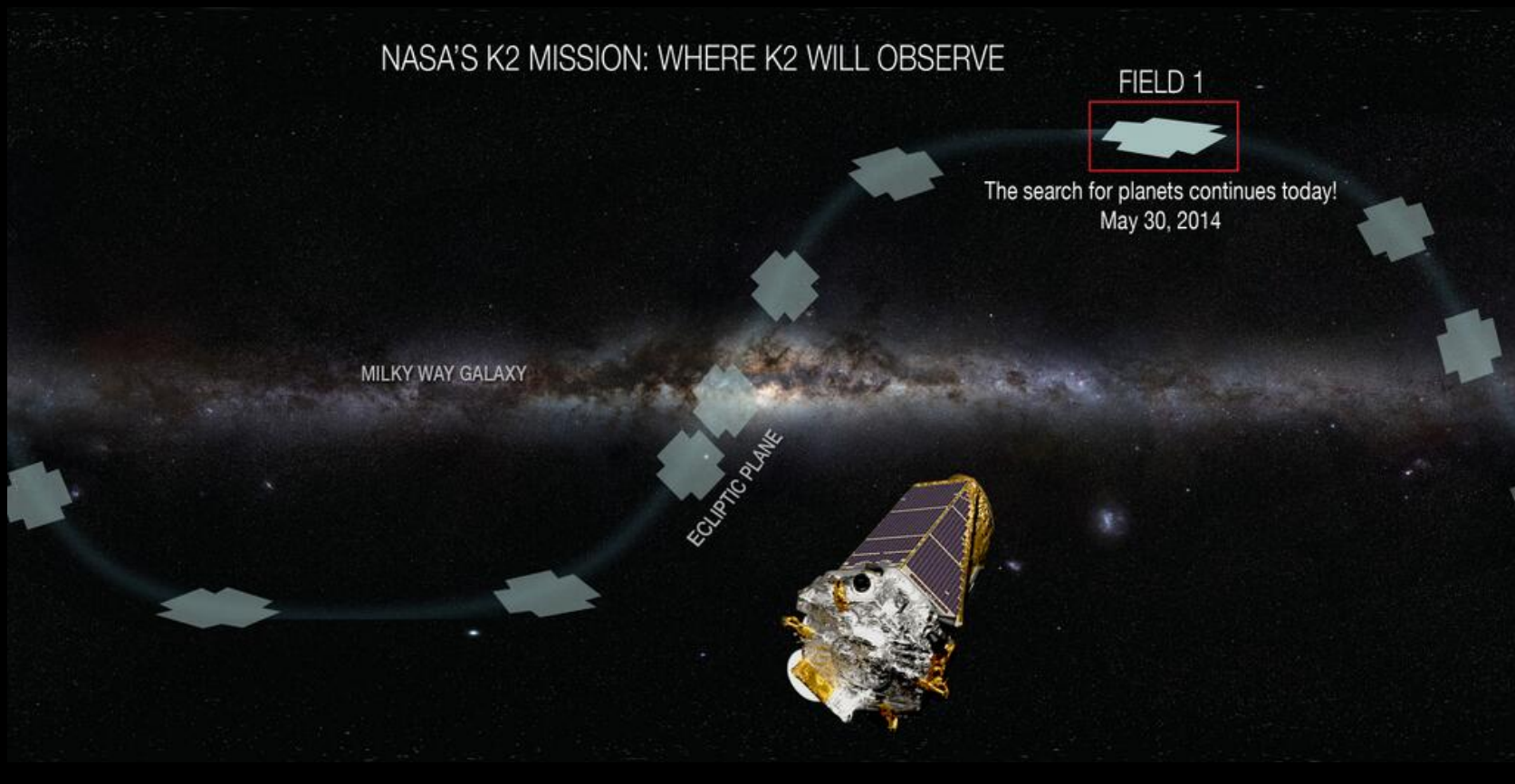
Ball Aerospace to the Rescue!



CONCEPTUAL ILLUSTRATION OF SPACECRAFT SOLAR DISTURBANCE. THE ACTUAL DISTURBANCE IS DUE TO PHOTON PRESSURE, NOT SOLAR WIND.

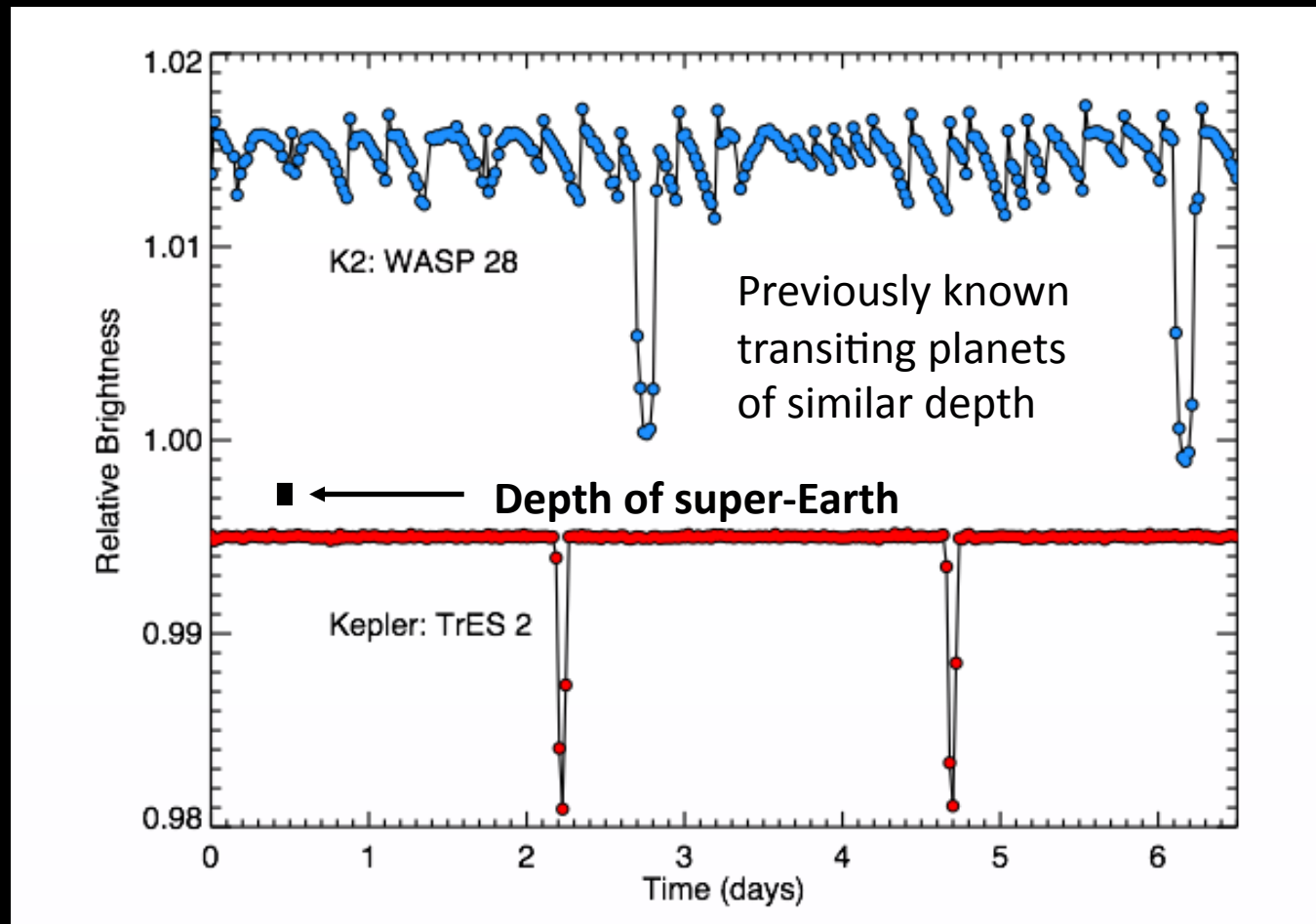
K2 Mission Profile

- Point at a field on the ecliptic for about 80 days
- Switch to a new field
- Entirely proposal driven: any astrophysics is fair game



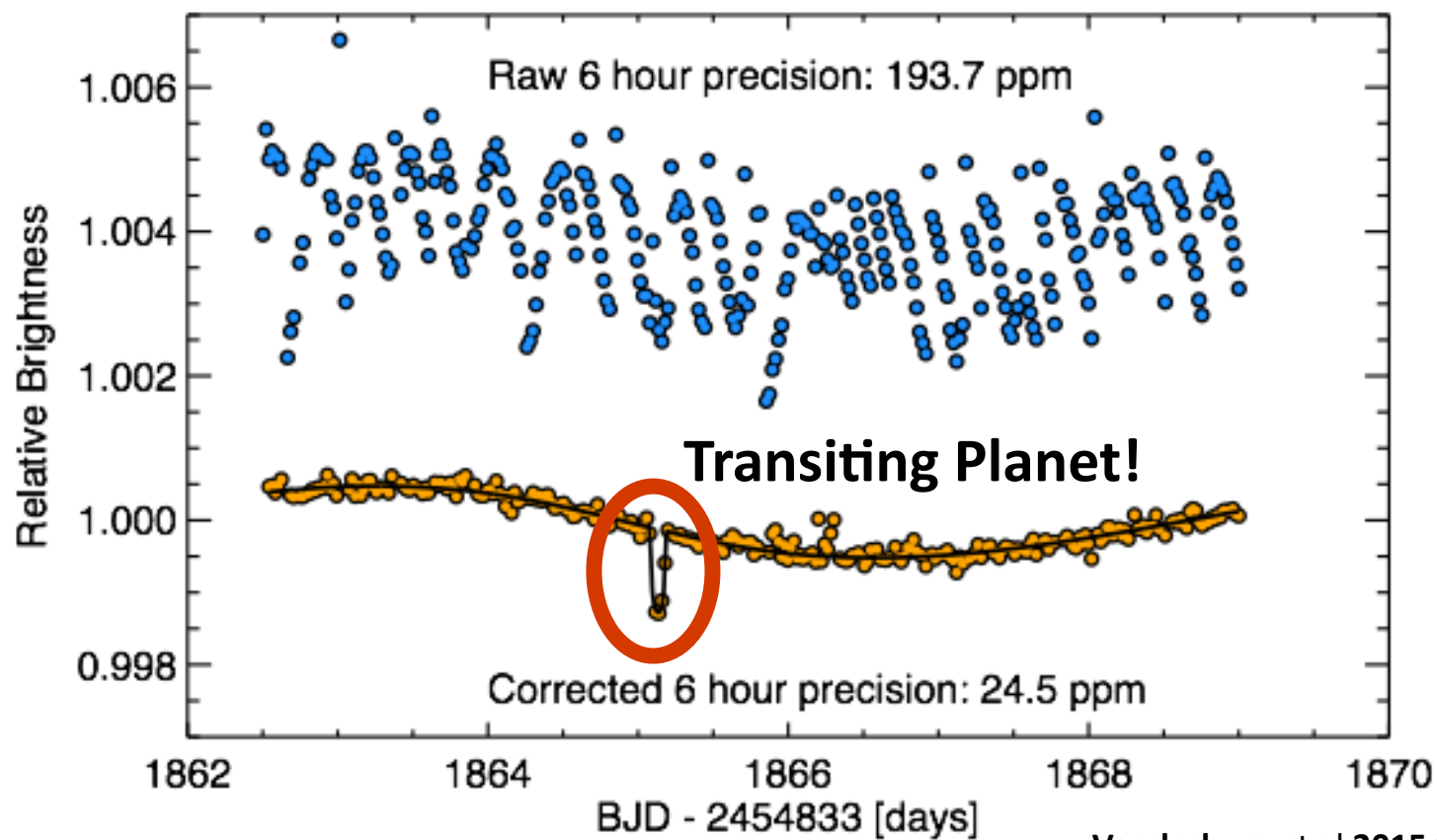
K2's Light Curves

- K2's small roll motion damages the light curves

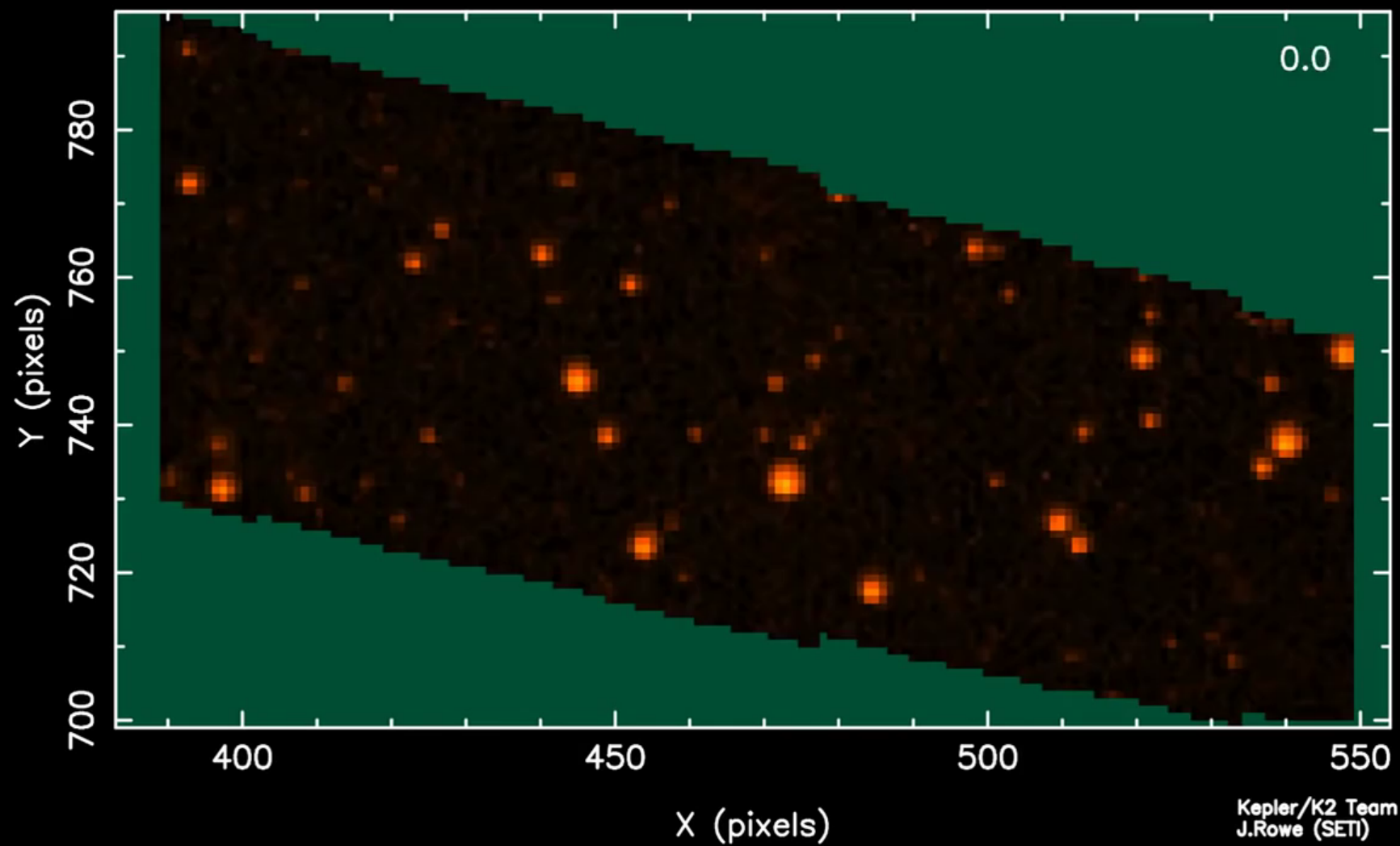


K2's Light Curves

- But they can be corrected: 39 confirmed K2 planets

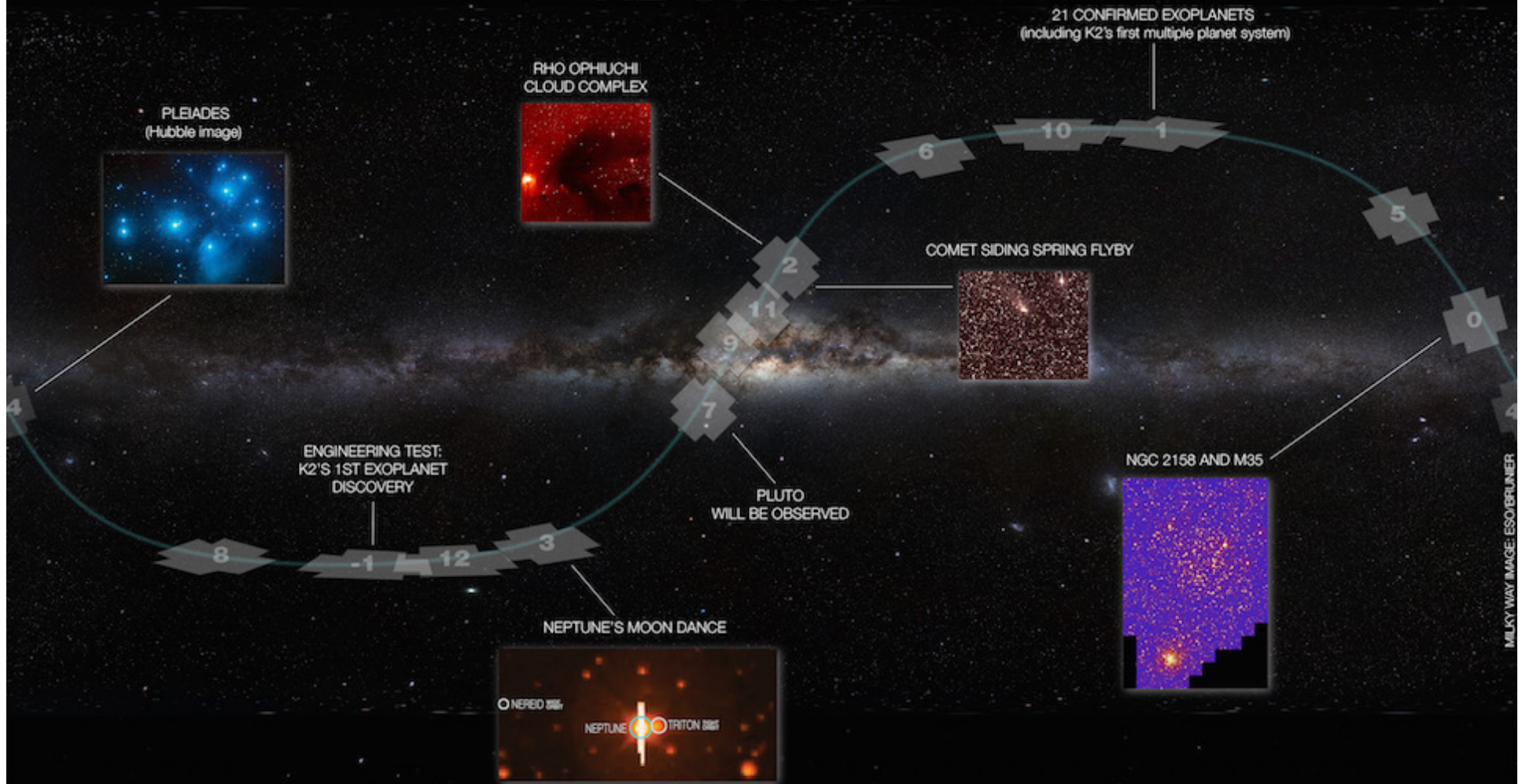


Neptune!





K2 Science



The Future: Doppler Method

- Doppler will get better

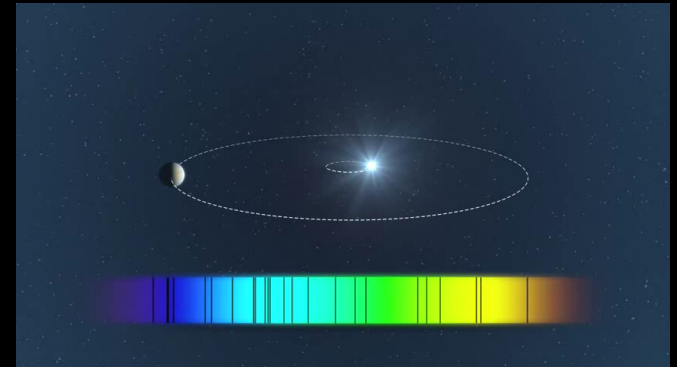
- HARPS North

- On La Palma

- Can detect speeds of 50 cm/sec

- Earth moves the Sun 9 cm/sec

- Doppler will maybe get to 10 cm/sec in a decade



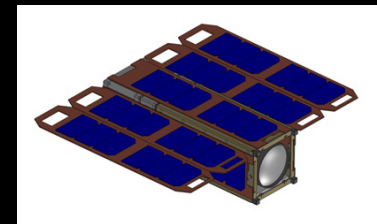
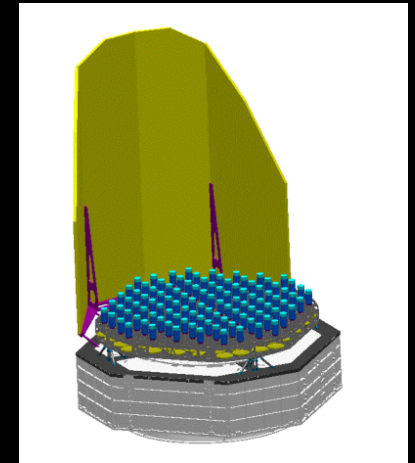
The Future: Transit Method

- Transiting Exoplanet Survey Satellite (TESS)
 - Selected for launch in 2017
 - Very high Earth orbit
 - Will survey the entire sky
 - 2 million brightest stars
 - Search the closest stars
 - Look at one star field for a few weeks, then move to another, and later return
 - Sensitive to Super Earths and larger
 - In short (2-month) orbits



The Future: Transit Method

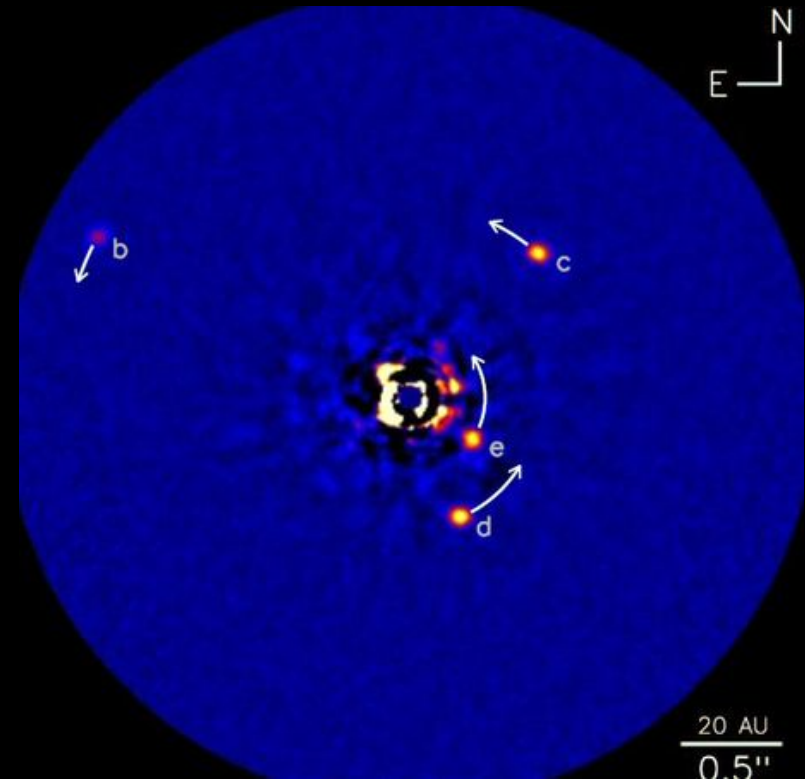
- Space missions on the Drawing Board
 - PLATO: European satellite with many cameras
 - Similar to TESS
 - ExoplanetSat: many small telescopes in space, each dedicated to a single star



- Several continuing Earth-based surveys

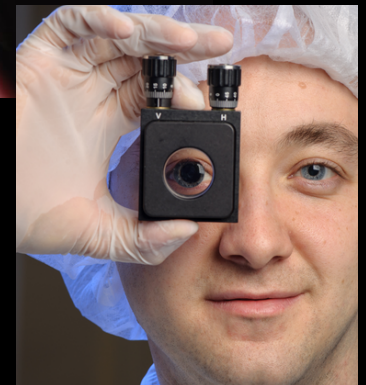
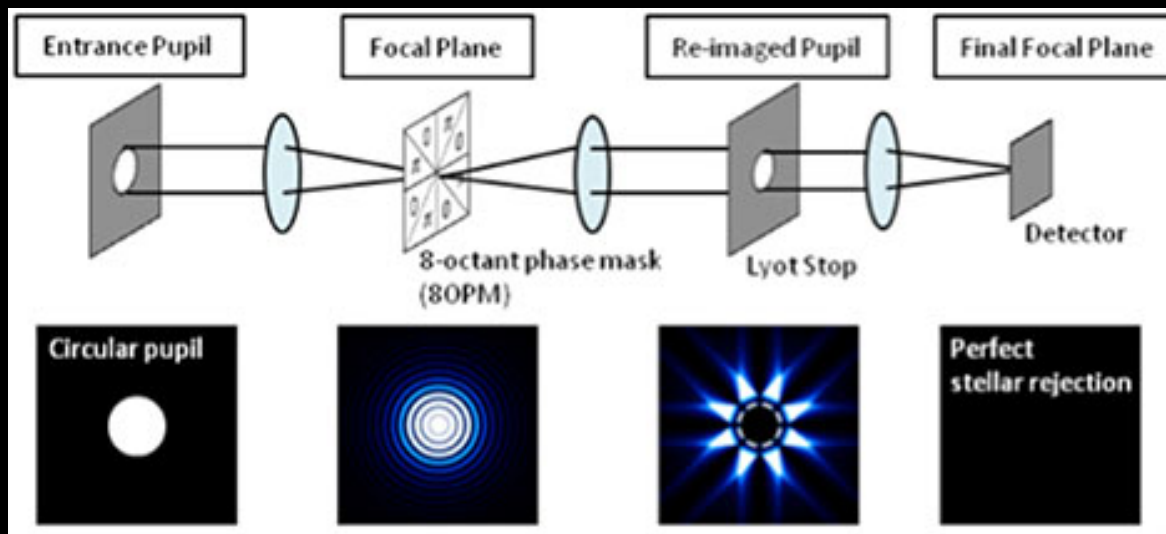
The Future: Direct Imaging

- Problem: block light from the star and not the planet
 - Done from the ground for Jupiter-size planets far from their star: HR 8799
 - Very difficult to get close to Earth-like distances from the star



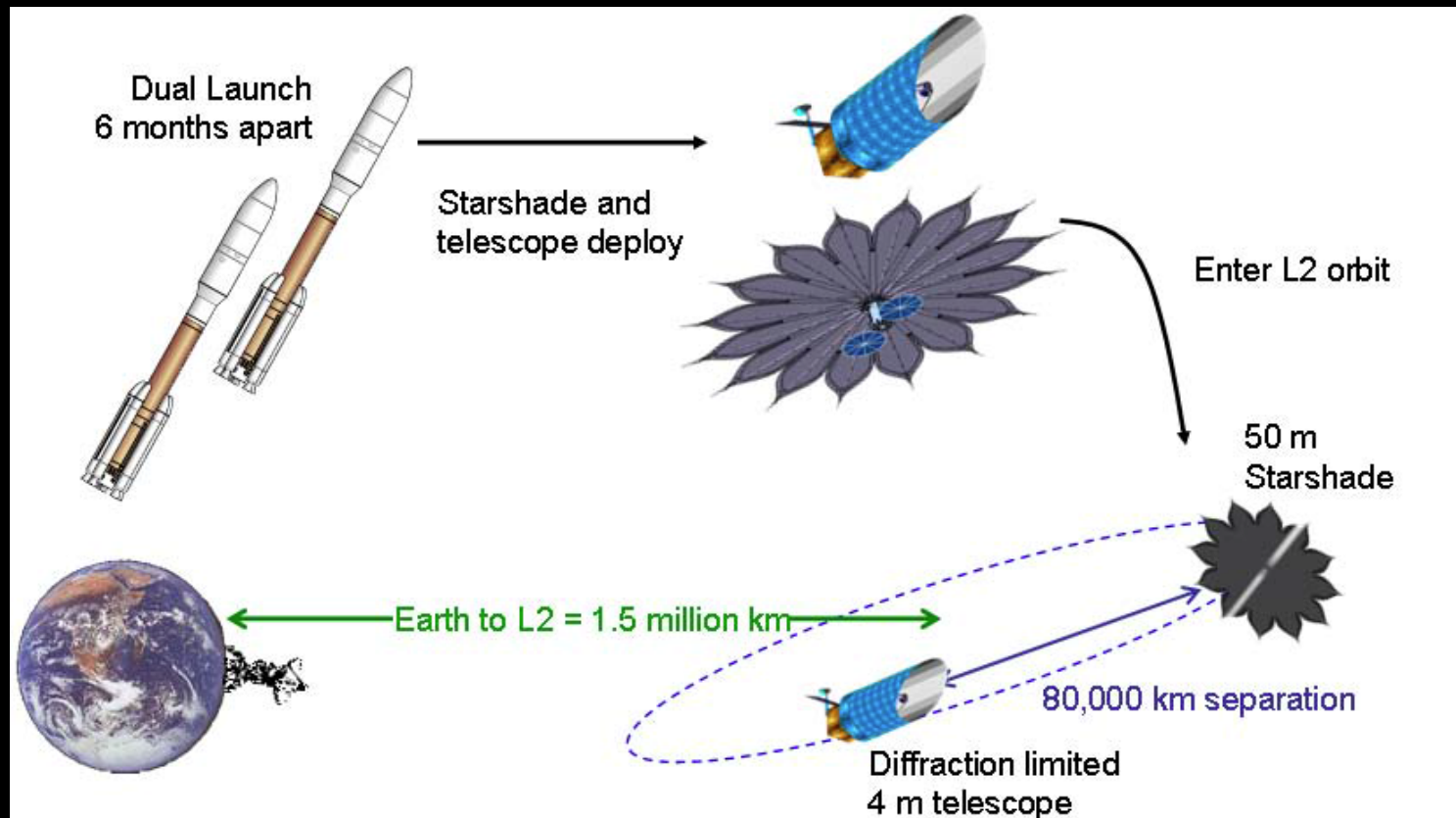
The Future: Direct Imaging

- Approach 1: block the star's light inside the telescope (coronagraph)
 - Difficult, requires extreme precision
 - Current status: Planned for the WFIRST telescope, to launch in the mid 2020's



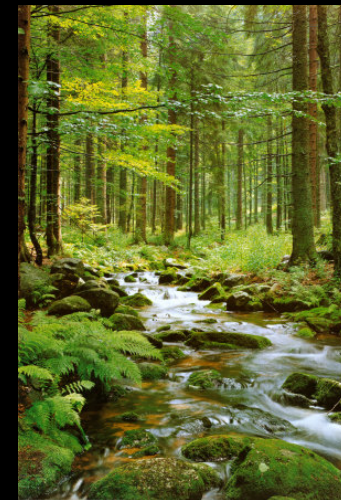
The Future: Direct Imaging

- Approach 2: Block the star's light with a starshade in front of the telescope



Direct Imaging: Difficult but Worth It!

- The best picture of a planet will be just a dot
- But from that dot we can measure its atmosphere
 - Oxygen would mean life!
 - Left by itself, oxygen is absorbed by the ground (rust)
 - The only way we know to replace oxygen is photosynthesis by living organisms



Conclusion

- The search for exoplanets takes us closer to answering
 - Why are we here?
 - Is the Earth common or rare?
 - Preliminary result: looks common!
 - Are we alone?
 - Is there life on other planets?
 - Don't know yet, but there are many places where life is possible
 - How did we get here?
 - How did our planet form?
 - Comparing with other planet systems helps us understand ours